# Search for a monochromatic component of solar axions using Fe-57

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# Axion



- Undiscovered pseudoscalar particle predicted to solve the ``strong CP'' problem
- $m_a$ ??,  $g_a$ ?? no prediction,

but  $m_a \propto g_a$ 

• One candidate of cold dark matter,

if  $m_a = 10^{-3} \sim 10^{-6} \text{ eV} \Rightarrow \text{Cavity experiments}$ 

 If m<sub>a</sub> (g<sub>a</sub>) is reasonably large, it is unlikely as CDM, but

 $\Rightarrow$  the sun can be a bright axion source

#### Solar axion





#### (off-topic) Search with $g_{a\gamma\gamma}$ Tokyo axion helioscope



- *BL*=4T×2m
- Improvements will be done in ~one year to explore m<sub>a</sub>~ eV region

# Monochromatic solar axion and its detection

- Like Mössbauer effect
- No need to move detector (~5eV width due to thermal motion)
- Not affected by the theoretical uncertainty of g<sub>avv</sub>





# Target region and other experimental constraints based on $g_{aNN}$

- Not so many experimental constraints
  - From J/ψ decay: m<sub>a</sub><~6 keV</li>
  - Similar solar <sup>57</sup>Fe axion search: m<sub>a</sub><745 eV (M. Krčmar *et al.*, PLB 442(1998)38)
  - SN1987A??? What is it?

Even if so, the hadronic axion window still exists arond  $m_{\rm a} \sim$  a few tens eV!!

- If *m*<sub>a</sub>~100 eV,
  - <sup>57</sup>Fe axion flux on the earth:  $\phi$ ~10<sup>12</sup> cm<sup>-2</sup> s<sup>-1</sup>
  - <sup>57</sup>Fe excitation rate: *R*~10<sup>2</sup> day<sup>-1</sup> g<sup>-1</sup>



#### **Detector setup I**

- <sup>57</sup>Fe foil and large area Si PIN photodiodes
- Cooled with dry ice (195 K)



### **Detector setup II**





- 2 × Si PIN photodiodes:
  - Hamamatsu S3584-06
  - Active area: 28mm×28mm
  - Thickness: 500µm
  - Specially packaged with low-BG ceramic
  - <sup>57</sup>Fe foil
  - Enriched: 95.85%

(natural abundance: 2.2%)

- 32mm × 32mm
- Thickness: (40±5)µm
  (att. length=20µm @14.4keV)
- Mass: 320mg

#### **Detector performance**



- Absorption in iron
- Acceptance
- Incident angle
- ⇒Estimated by Geant4 based MC

>14.8%

for 14.4keV from the foil





### Measurement

- From July 26<sup>th</sup>, 2005 to September 3<sup>rd</sup>, 2005
  - Two types of foils were attached



- Background subtraction corresponds to signals from 93.6% <sup>57</sup>Fe
- During all measurements,
  - Same apparatus were used except for the foil
  - Stable operation
    - Temperature of Holder (198±1) K
    - Temperature of 1<sup>st</sup> stage FET (212±1) K
    - Gain fluctuation < ±0.3% (Checked by 59.5keV from <sup>241</sup>Am)



#### **Obtained spectrum**







#### **Results**



 From the obtained 95% limit: < 7.93×10<sup>-5</sup> cps ±2σ region: 95.45%
 <sup>57</sup>Fe mass difference: 197mg Detection efficiency: 14.8% Branch of γ-ray emission: 10.5%
 ⇒ Excitation rate of <sup>57</sup>Fe: R < 2.35×10<sup>3</sup> /day/g (95% C.L.)

(Factor 144 improvement)

- This corresponds to:
  - *f<sub>a</sub>* > 2.89×10<sup>4</sup> GeV, *m<sub>a</sub>*< 216 eV (if *z*=0.56, S=0.5) (Factor 3.5 improvement)

### **Future prospects**

Towards the hadronic axion window

$$(m_a \sim a \text{ few tens eV})$$

- Development of <u>a new detector made of iron</u> is necessary
- If iron detector is possible,
  - $\Rightarrow$  2 order improvements of S/N, even if the same BG level
- One candidate:

New generation semiconductor:  $\beta$ -FeSi<sub>2</sub> (iron silicide)

 Possibility for the application as a particle detector is under investigating (e.g. growth of bulk)





诱明窓

無反射膜 n-Si

n-β-FeSi2 p-β-FeSi2

センサ

# Summary



- A new search for solar axion was performed
- This search is based on  $g_{\rm aNN}$
- <sup>57</sup>Fe foil and silicon PIN photodiodes were used
- From the absence of <sup>57</sup>Fe related signals, a new experimental constraint was obtained
  m<sub>a</sub> < 216 eV, f<sub>a</sub> > 2.89×10<sup>4</sup> GeV
  (z=0.56, S=0.5)