

# Recent progress in positronium laser cooling

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N. Oshima<sup>6</sup>, B.E. O'Rourke<sup>6</sup>, K. Michishio<sup>6</sup>, K. Ito<sup>6</sup>, K. Kumagai<sup>6</sup>, R. Suzuki<sup>6</sup>,  
S. Fujino<sup>7</sup>, T. Hyodo<sup>8</sup>, I. Mochizuki<sup>8</sup>, K. Wada<sup>8</sup>, T. Kai<sup>9</sup>



[https://tabletop.icepp.s.u-tokyo.ac.jp/?page\\_id=365](https://tabletop.icepp.s.u-tokyo.ac.jp/?page_id=365)

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# Acknowledgement



創発の研究支援事業

*Fusion Oriented Research for disruptive Science and Technology*

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## 1. Motivation for Ps laser cooling

- Physics target: Bose-Einstein condensate (BEC), precision spectroscopy, antimatter gravity
- Fast Ps cooling by a combination of thermalization and laser cooling

## 2. Development of Ps cooling laser

- Requirements
- Chirped Pulse-Train Generator (CPTG)

## 3. Experiment at KEK slow positron facility (SPF)

- Experimental setup
- Current status of proof-of-principle experiment
- Prospects

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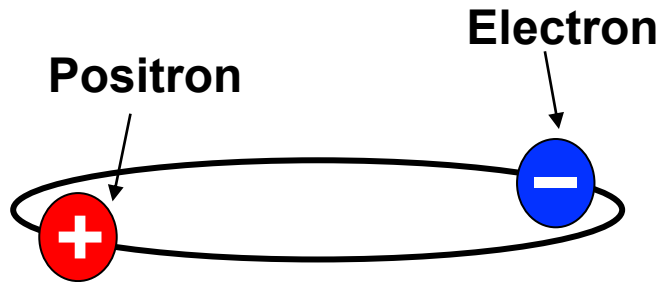
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# Positronium (Ps) is a good probe for fundamental physics

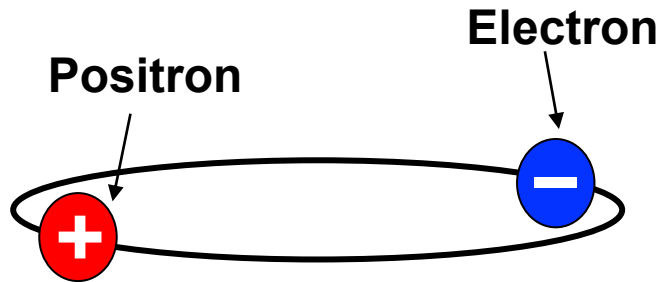


Bound state of an electron ( $e^-$ ) and a positron ( $e^+$ )

The Lightest and Exotic Atom

- ✓ Exotic atom with antiparticle
  - Good to explore the mystery of antimatter
- ✓ Purely leptonic system
  - Experiments and theoretical calculations can be compared in high precision without uncertainties of hadronic interactions.

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The Lightest and Exotic Atom

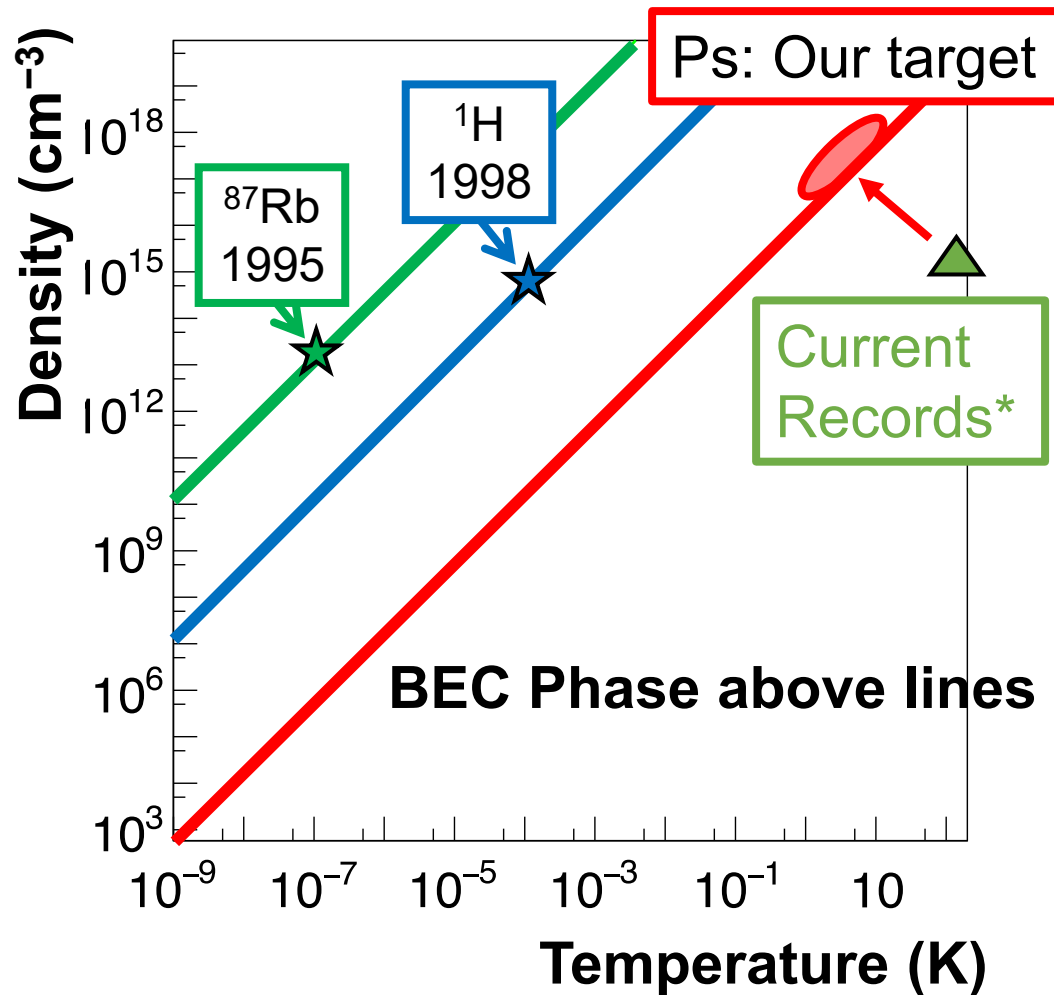
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  - Good to explore the mystery of antimatter
- ✓ Purely leptonic system
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In this talk, we focus on *ortho*-positronium (*o*-Ps), the spin-triplet state with a lifetime of 142 ns.

*o*-Ps has much longer lifetime than the spin-singlet *para*-positronium (*p*-Ps), which has a lifetime of 125 ps.

## Physics Target (1):

# Positronium Bose-Einstein condensate (Ps-BEC)



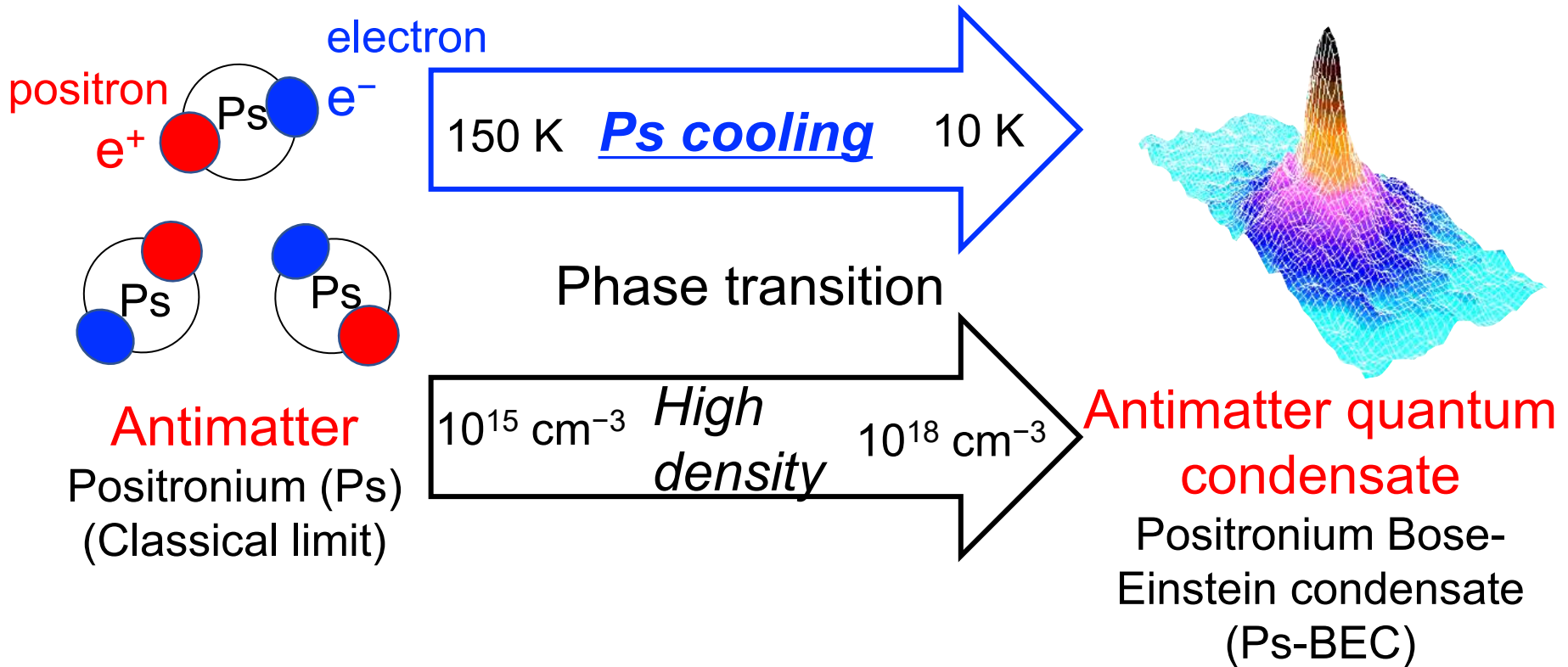
- Ps must be **dense** and **cold**
- High critical temperature because of Ps light mass (14 K at  $10^{18} \text{ cm}^{-3}$ )
- One of the best candidates for the **first antimatter BEC**
- BEC is “Atomic laser”. We would like to make the **first antimatter laser** and perform new experiments using the coherency of Ps-BEC.

\* : S. Mariuzzi *et al.*, Phys. Rev. Lett. **104**, 243401 (2010)

\* : D. Cassidy *et al.*, physica status solidi **4**, 3419 (2007)

# Physics Target (1): Ps-BEC

Realization of Ps-BEC at:  
**Low temperature (10 K)** and  
quite high density ( $10^{18} \text{ cm}^{-3}$ )

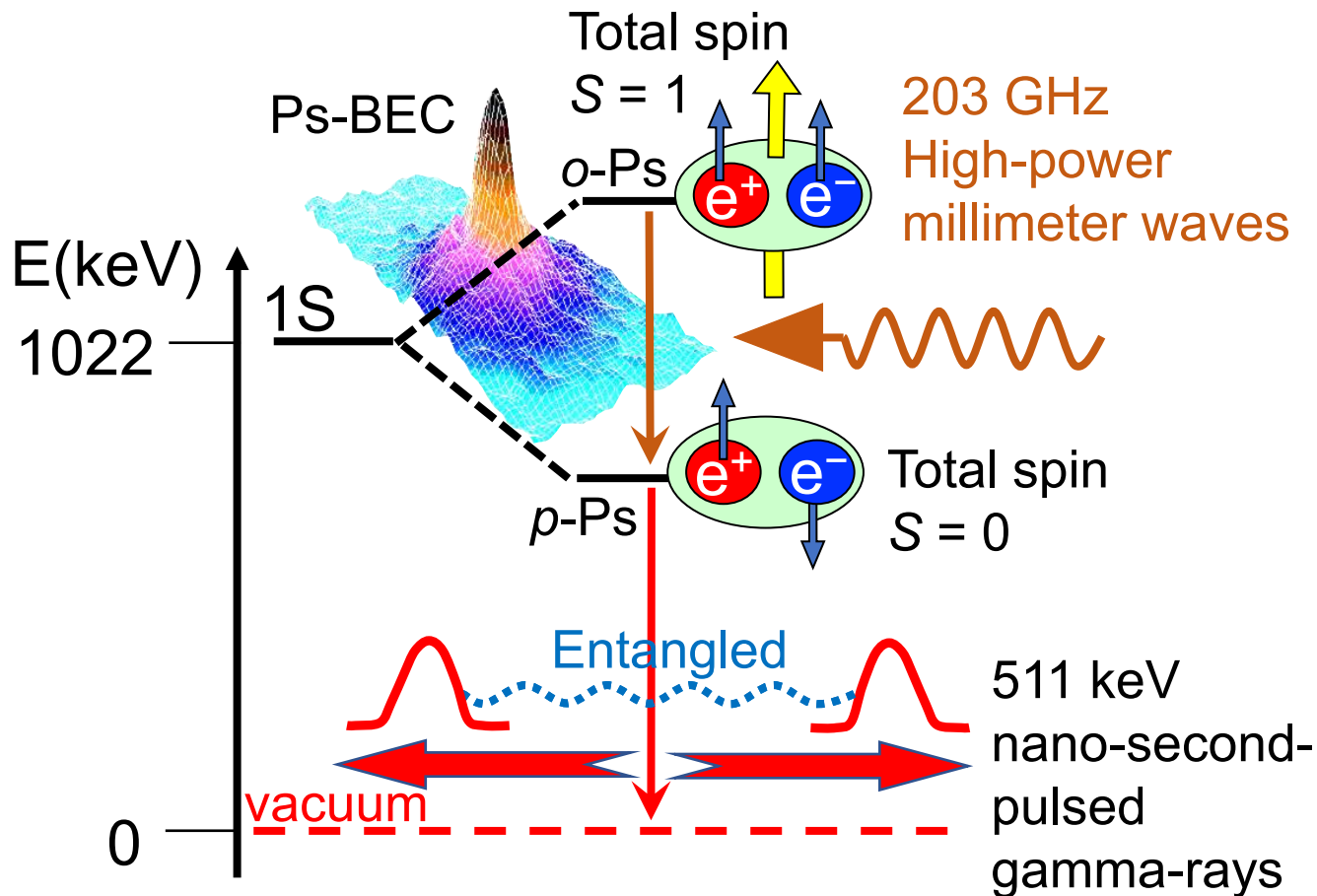


K. Shu *et al.*, J. Phys. B **49**, 104001 (2016), A. Ishida *et al.*, JJAP Conf. Proc. **7**, 011001 (2018).



# Physics Target (1): Ps-BEC

Self-annihilations of Ps-BEC can generate  
2 coherent and entangled gamma-rays:  
Realization of **gamma-ray lasers**

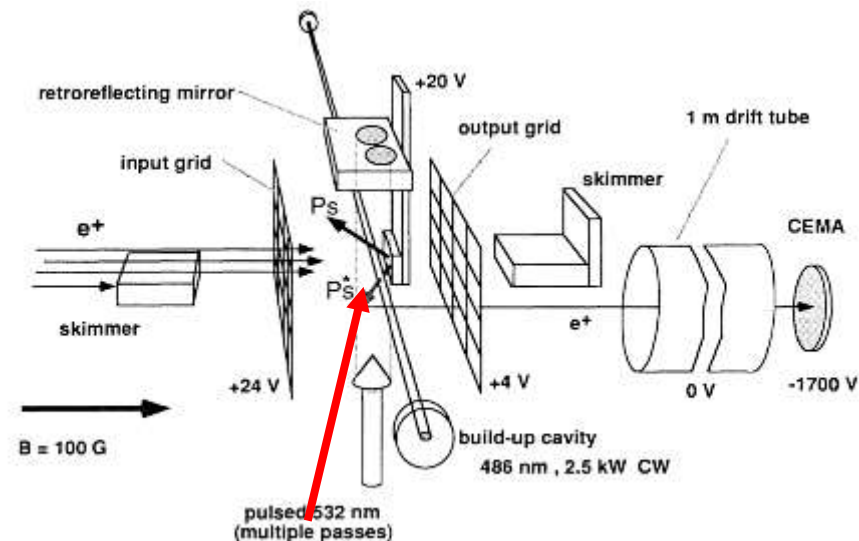
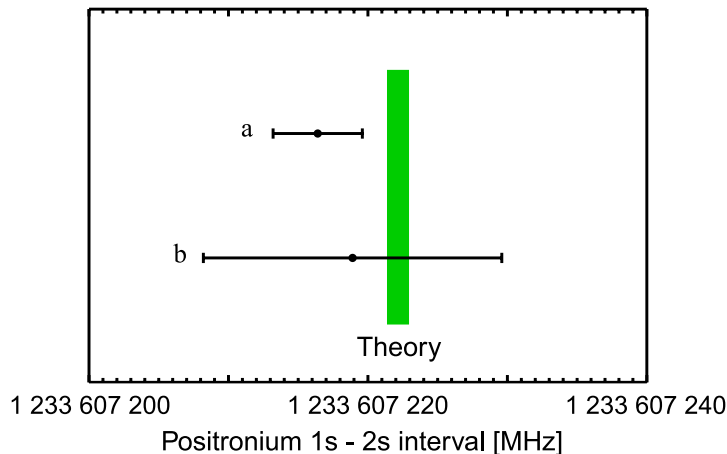


# Physics Target (2): Precision spectroscopy

$m_{\text{Ps}} \approx 2m_e$ : The lightest atom  $\rightarrow$  large velocity

1S—2S precision spectroscopy:  $1\,233\,607\,216.4 \pm 3.2$  MHz (2.6 ppb)  
Ps cooled down to 10 K can improve the precision by an order of magnitude

*Savely G. Karshenboim / Physics Reports 422 (2005) 1–63*



Ps at around 600 K

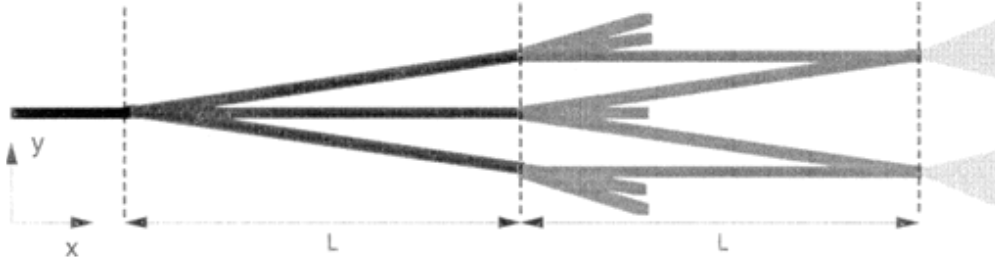
M.S. Fee *et al.*, Phys. Rev. Lett. **70**, 1397 (1993).

The high Ps temperature is the largest source of the uncertainty and bias.

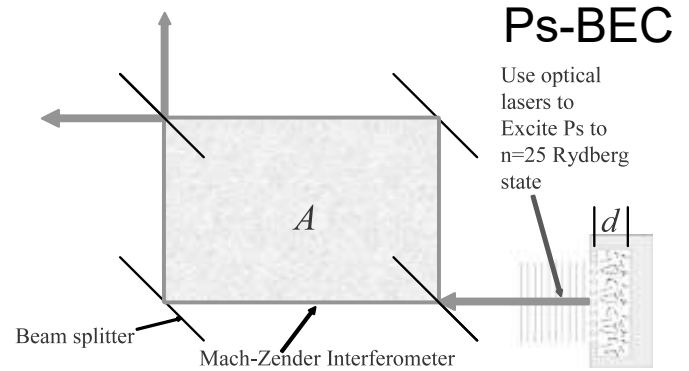
# Physics Target (3): Antimatter gravity

## Interferometry

coherence will be increased  
by cooling Ps



T. J. Phillips, *Hyperfine Interactions* **109**, 357 (1997).

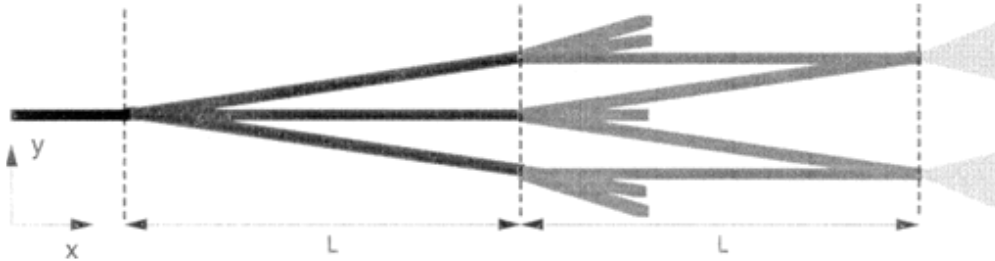


D. B. Cassidy and A. P. Mills, Jr.,  
*phys. stat. sol. (c)* **4**, 3419 (2007).

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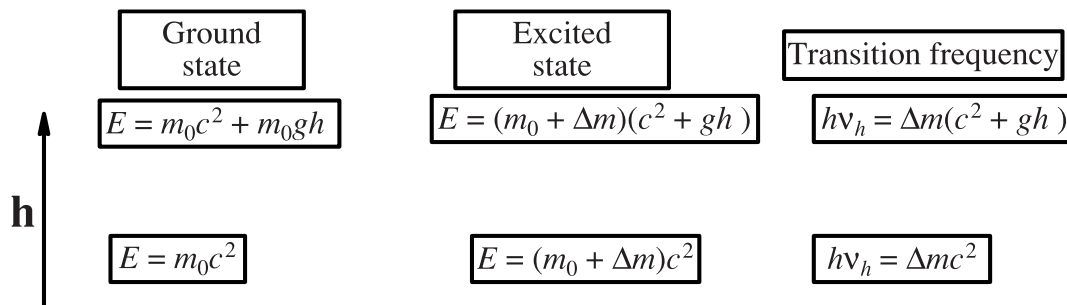
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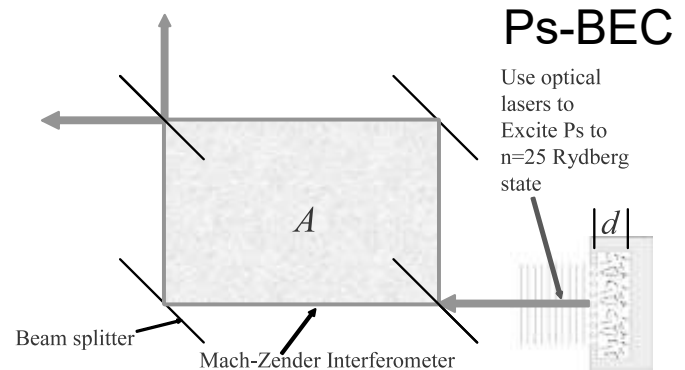
T. J. Phillips, *Hyperfine Interactions* **109**, 357 (1997).

## Spectroscopy



**Fig. 1.** Derivation of the gravitational red shift.

S. G. Karshenboim, *Astron. Lett.* **35**, 663 (2009).



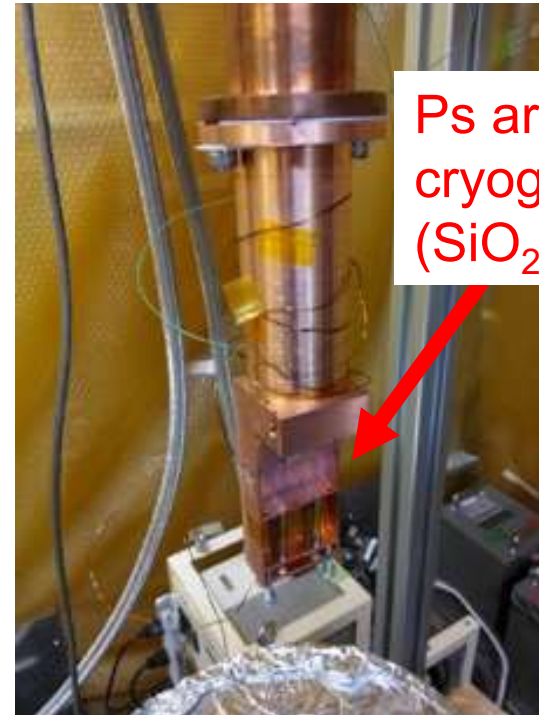
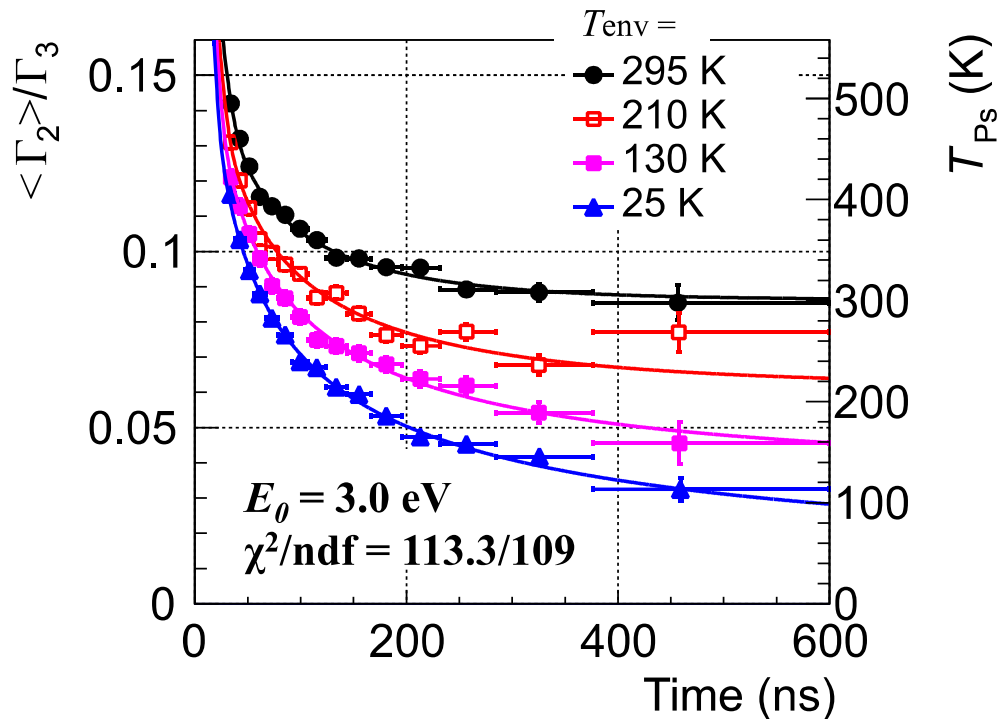
D. B. Cassidy and A. P. Mills, Jr.,  
*phys. stat. sol. (c)* **4**, 3419 (2007).

Different distance  
from the Sun:

$$\frac{\Delta U(r_{\max}) - \Delta U(r_{\min})}{c^2} \simeq 3.2 \times 10^{-10}.$$

(0.32 ppb)  
cf. Ps 1S—2S : 2.6 ppb

**Thermalization cooling is too slow to get 10 K.**  
**→ We need faster cooling: Ps laser cooling.**



PHYSICAL REVIEW A **104**, L050801 (2021)

Letter

### Observation of orthopositronium thermalization in silica aerogel at cryogenic temperatures

Kenji Shu<sup>\*</sup>, Akira Ishida<sup>†</sup>, Toshio Namba, and Shoji Asai

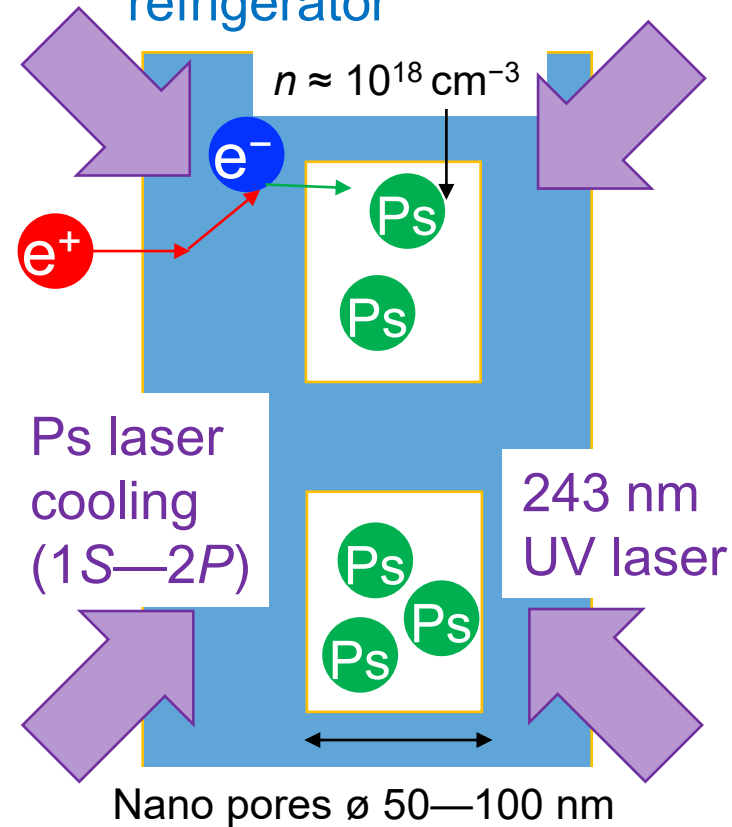
*Department of Physics, Graduate School of Science, and International Center for Elementary Particle Physics,  
 The University of Tokyo, 7-3-1 Hongo, Bunkyo-ku, Tokyo 113-0033, Japan*

Nagayasu Oshima, Brian E. O'Rourke, and Kenji Ito

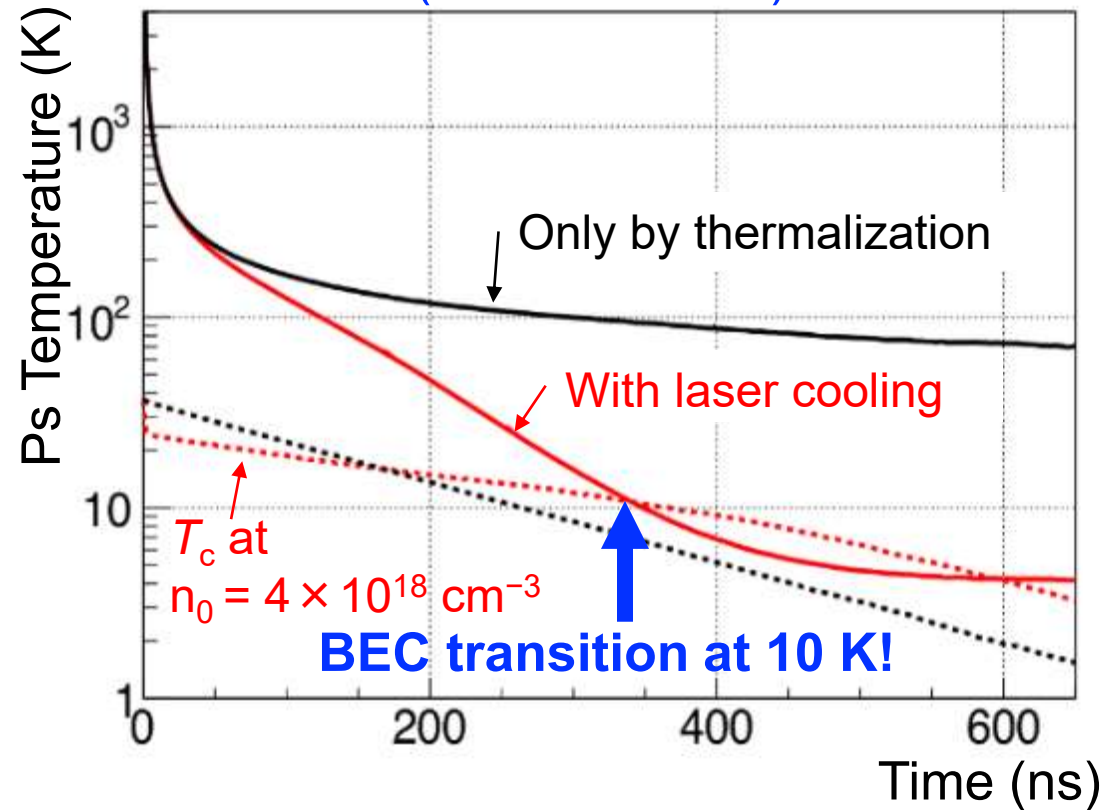
*National Institute of Advanced Industrial Science and Technology (AIST), Tsukuba-Central 2,  
 1-1-1 Umezono, Tsukuba, Ibaraki 305-8568, Japan*

# Combination of **Thermalization** and **Laser cooling** is suitable for fast Ps cooling to realize Ps-BEC.

Silica ( $\text{SiO}_2$ ) cooled down by cryogenic refrigerator



Ps temperature evolution (MC simulation)



K. Shu *et al.*, J. Phys. B **49**, 104001 (2016), A. Ishida *et al.*, JJAP Conf. Proc. **7**, 011001 (2018).

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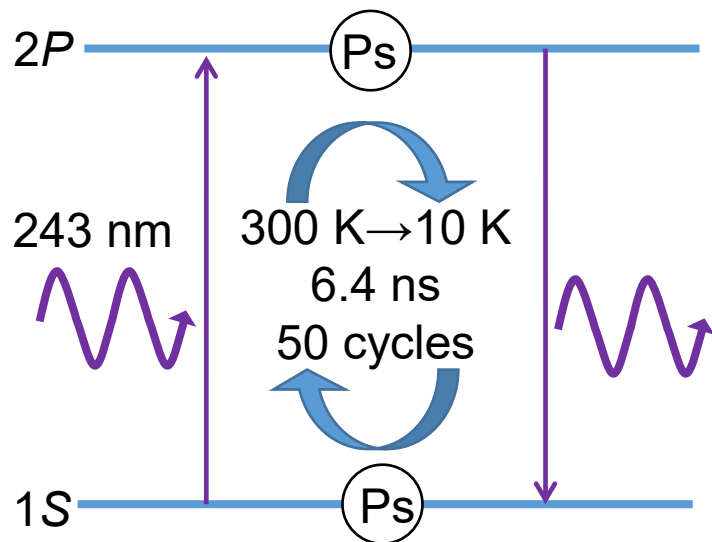
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# Requirements for Ps cooling laser

1. Long pulse duration and sufficient pulse energy

← Ps lifetime: 142 ns



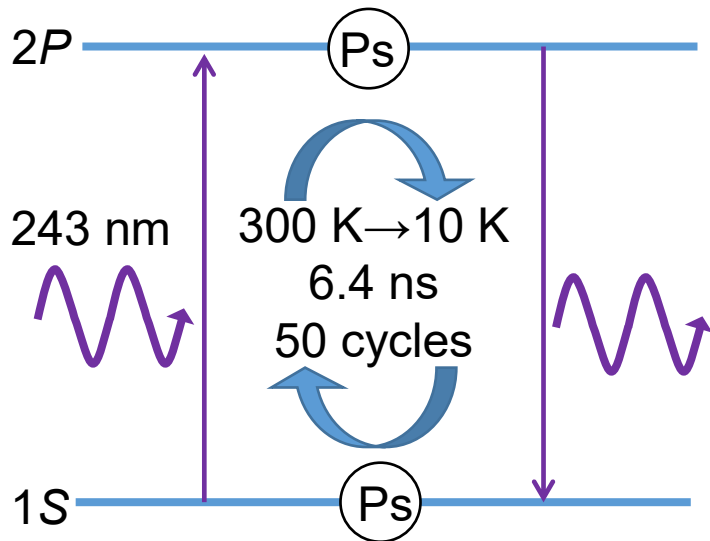
- $1S \rightarrow 2P$  (243 nm)
- $6.4 \text{ ns} \times 50 \approx 300 \text{ ns} \rightarrow$   
Cool down Ps with 300-ns  
single pulse
- Pulse energy  $\geq 40 \mu\text{J}$



# Requirements for Ps cooling laser

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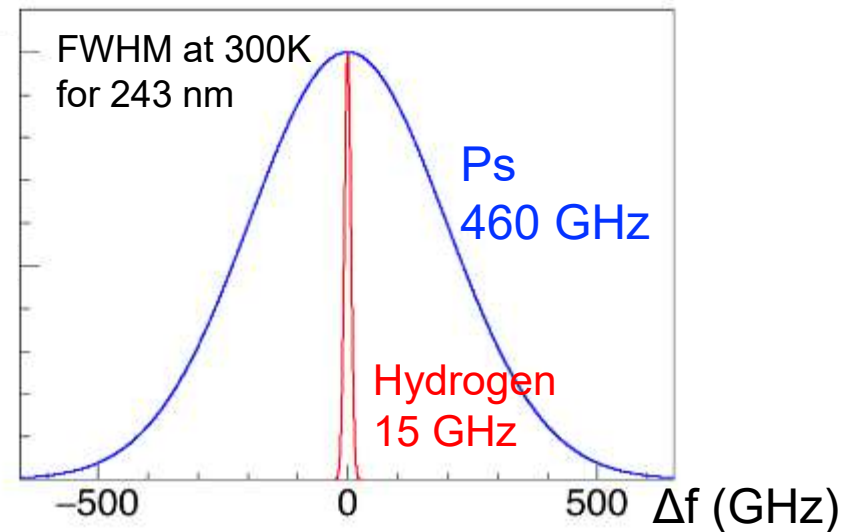
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Cool down Ps with 300-ns single pulse
- Pulse energy  $\geq 40 \mu\text{J}$

2. Broadband and frequency-chirped

← Ps light mass:  $2m_e$



- Doppler broadening is 30 times larger than hydrogen. To follow the change in the Doppler profile by cooling (300 to 10 K in 300 ns)  $\rightarrow$  Broadband ( $\geq 10 \text{ GHz}$ ) and Frequency chirp  $\geq +0.2 \text{ GHz ns}^{-1}$

**$\rightarrow$  243 nm sub- $\mu\text{s}$  pulsed, broadband, and frequency-chirped laser**

# Chirped Pulse-Train Generator (CPTG)

PHYSICAL REVIEW APPLIED 16, 014009 (2021)

## Theoretical Analysis and Experimental Demonstration of a Chirped Pulse-Train Generator and its Potential for Efficient Cooling of Positronium

K. Yamada<sup>1</sup>, Y. Tajima<sup>2</sup>, T. Murayoshi<sup>1</sup>, X. Fan<sup>1,†</sup>, A. Ishida<sup>1</sup>, T. Namba<sup>3</sup>, S. Asai<sup>1</sup>,  
M. Kuwata-Gonokami<sup>1</sup>, E. Chae<sup>2,4,5</sup>, K. Shu<sup>1,2,4</sup> and K. Yoshioka<sup>1,2,4,\*</sup>

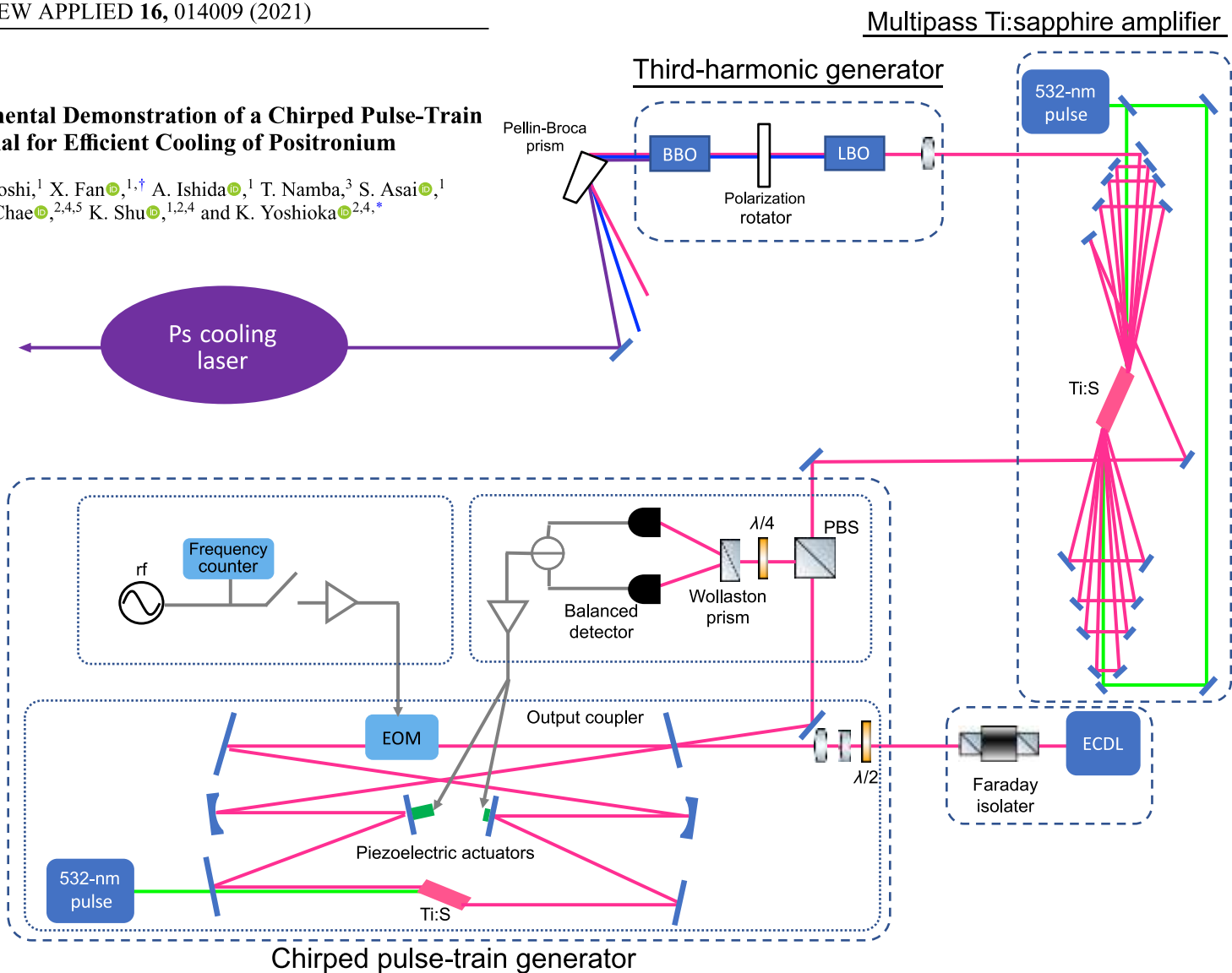
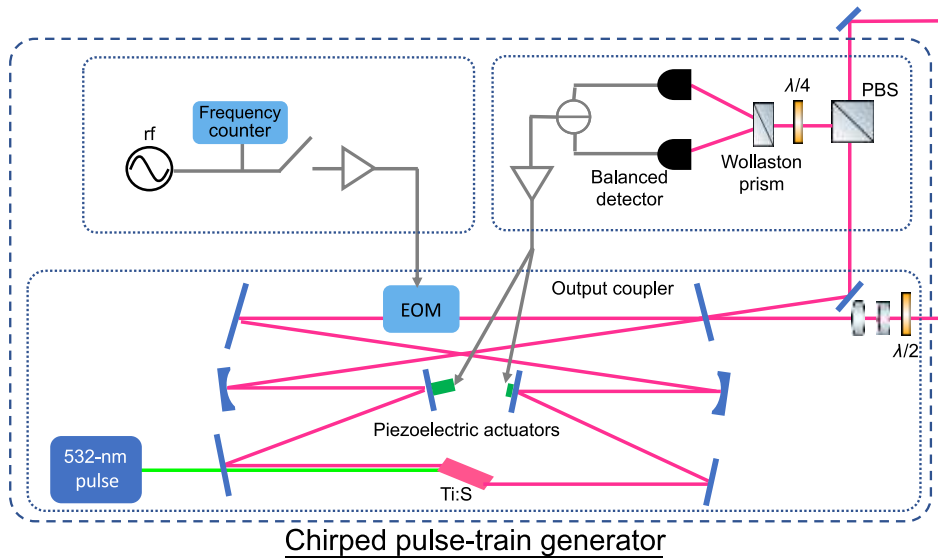
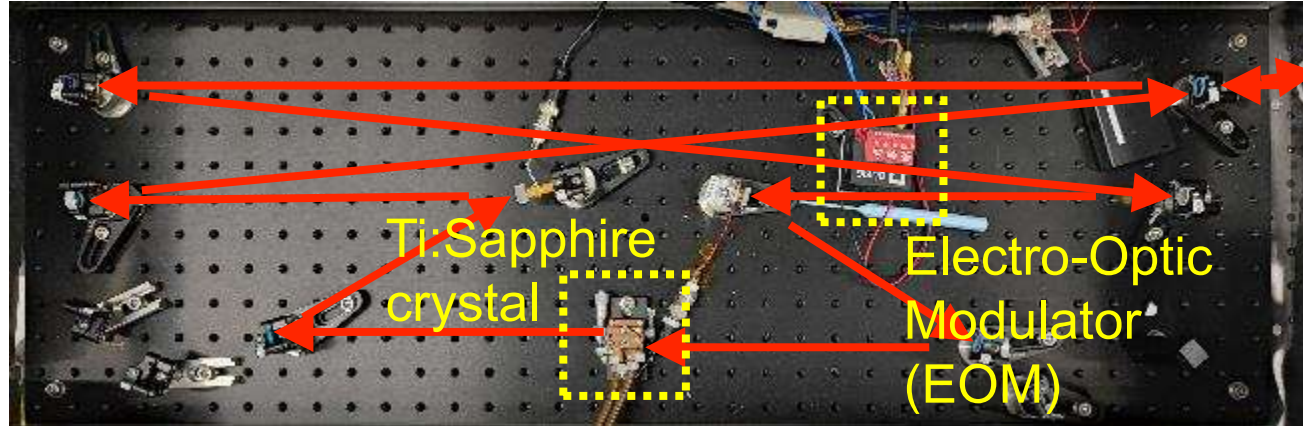
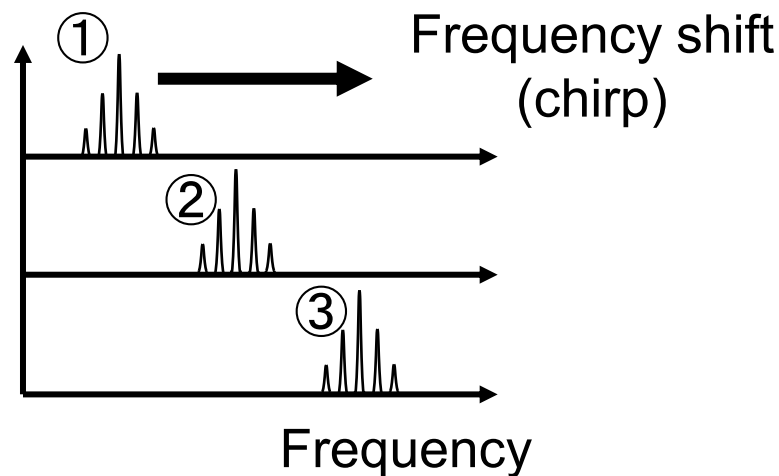
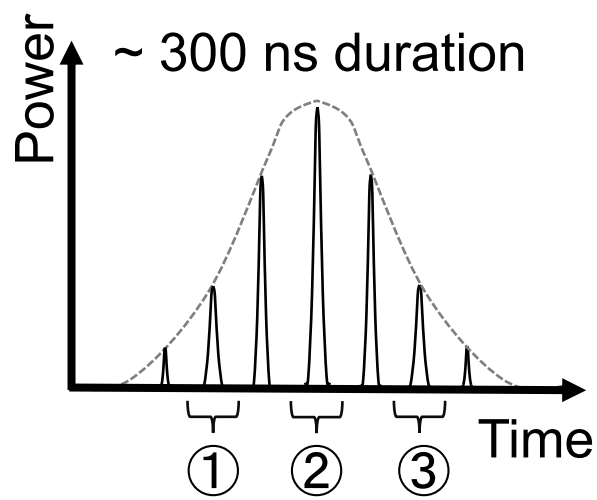
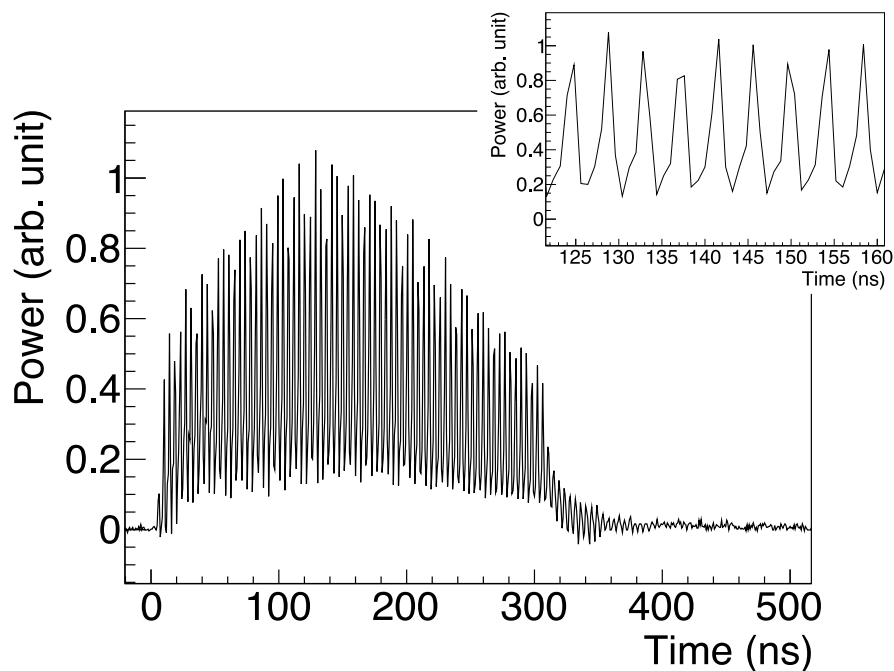


FIG. 1. Schematic of the prototypical cooling laser system. ECDL, external-cavity diode laser;  $\lambda/2$ , half-wave plate;  $\lambda/4$ , quarter-wave plate; EOM, electro-optic phase modulator; rf, driving radio frequency; Ti:S, Ti:sapphire crystal; PBS, polarizing beam splitter; LBO, lithium triborate crystal; BBO,  $\beta$ -barium borate crystal.

# Ps cooling laser



# Achieved 300-ns pulse duration



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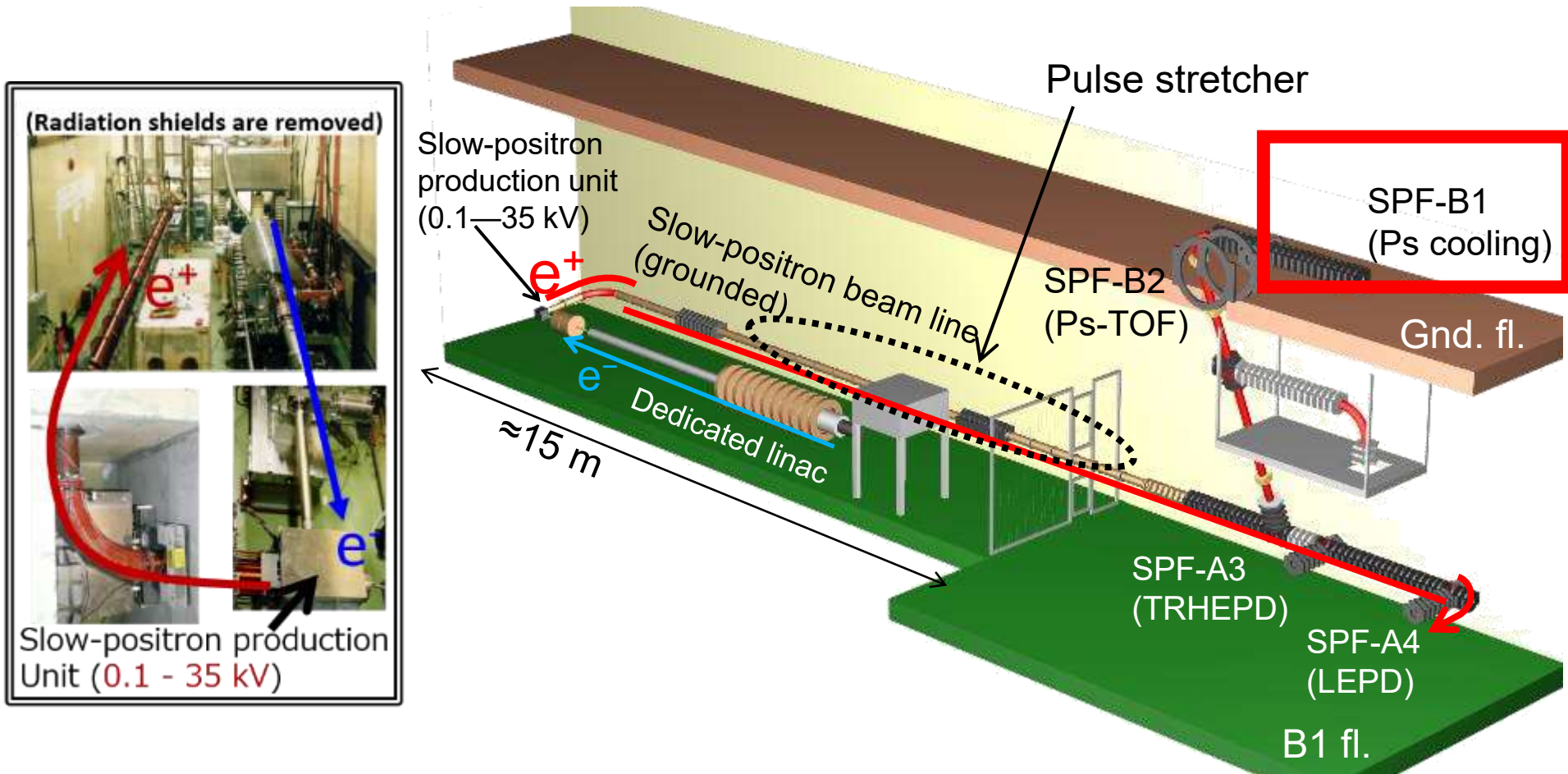
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We are trying a proof-of-principle experiment to **laser-cool Ps in vacuum** at **KEK-SPF (Slow Positron Facility)**, Tsukuba, Japan.

- Highest intensity ( $5 \times 10^7$  slow  $e^+$   $s^{-1}$ )
- Variable energy (0.1—35 keV)

Samples and detectors are electrically grounded.





# Ps Doppler cooling experiment with 3 lasers

## —1 cooling (243 nm) and 2 probe (243 & 532 nm) lasers—

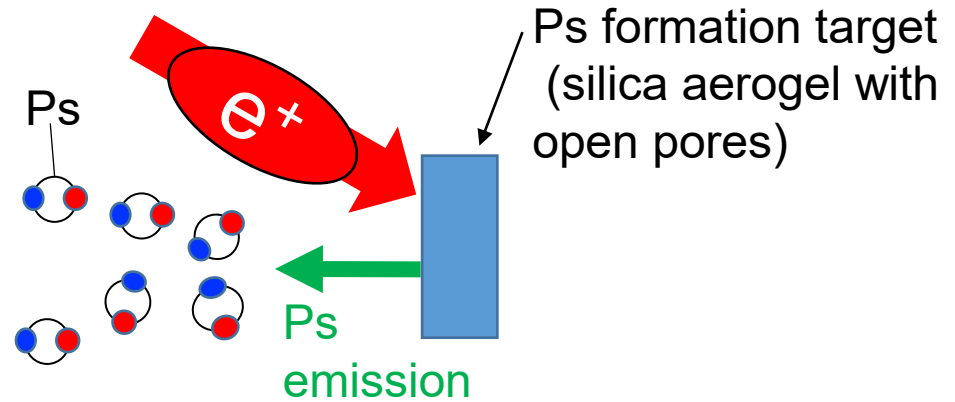
### (a) Laser irradiation of Ps

Cooling by laser **A**,  
Probe by lasers **B** & **C**

**A** : 243 nm (cooling)

**B** : 243 nm (probe)

**C** : 532 nm (probe)






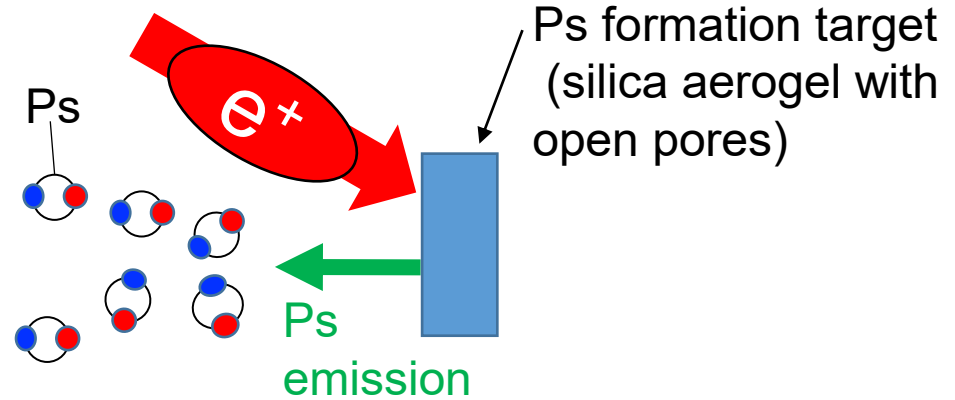
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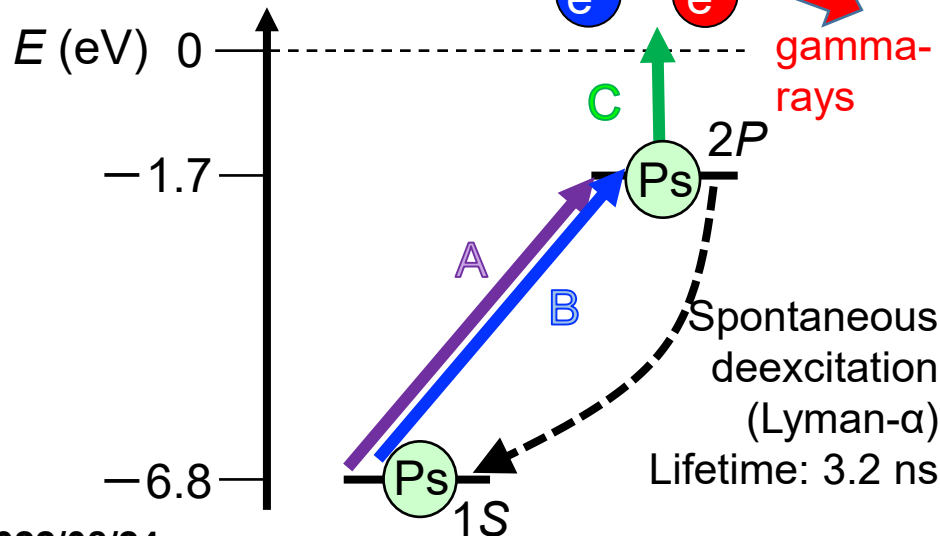
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**B** : 243 nm (probe)   
**C** : 532 nm (probe) 



### (b) Ps cooling cycle

$1S \rightarrow 2P \rightarrow 1S \rightarrow 2P \dots$ ,  
followed by  
 $1S \rightarrow 2P \rightarrow \text{ionization to detect Ps in } 2P \text{ state.}$








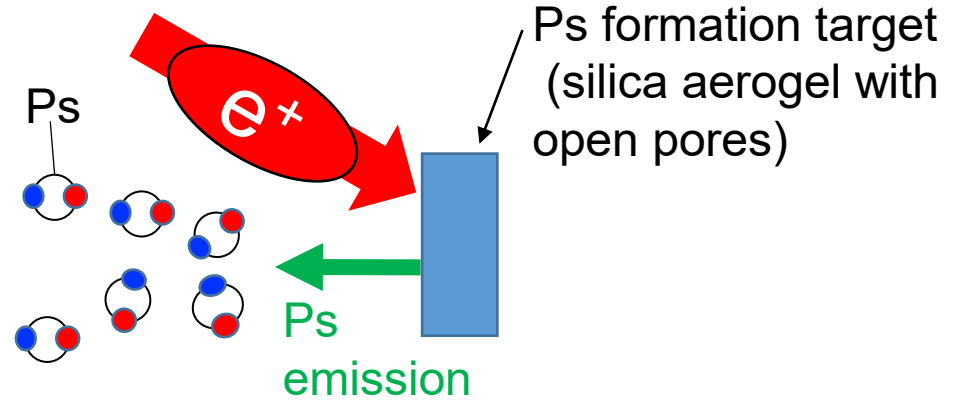
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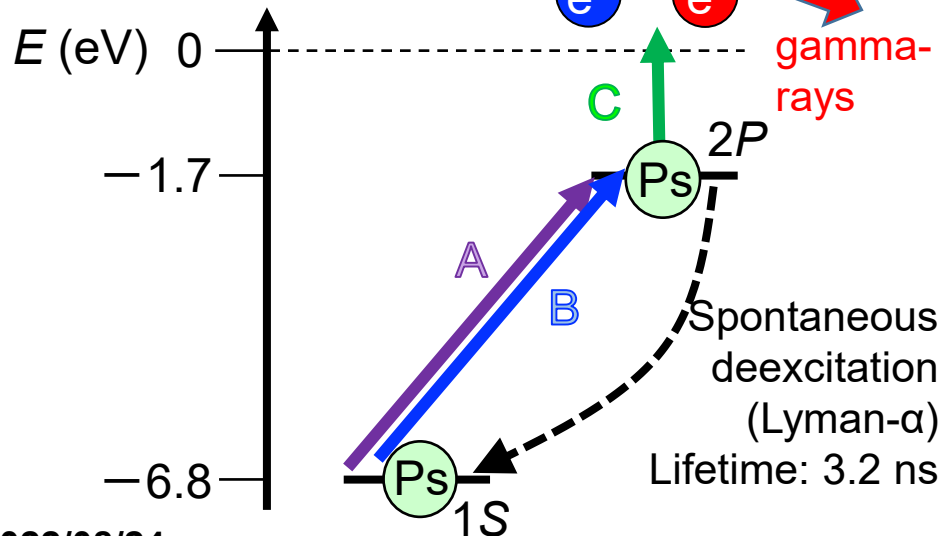
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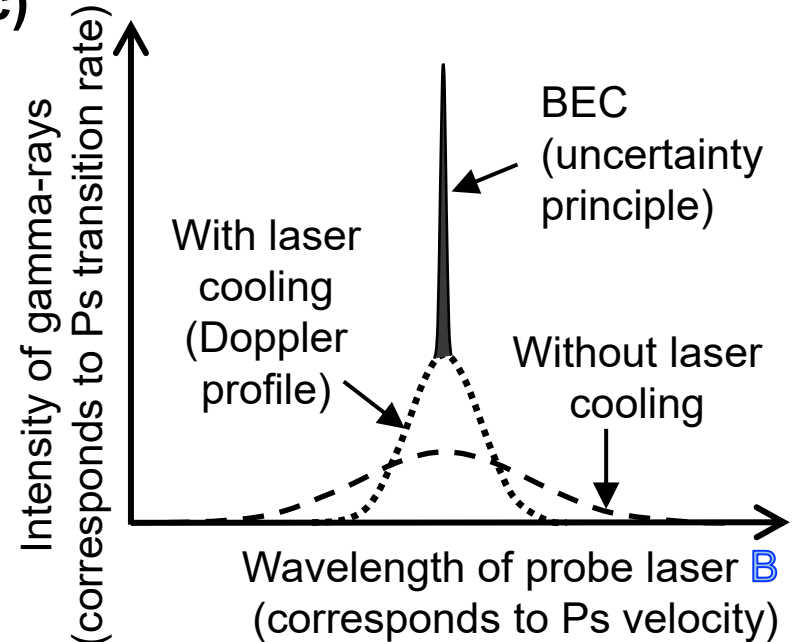


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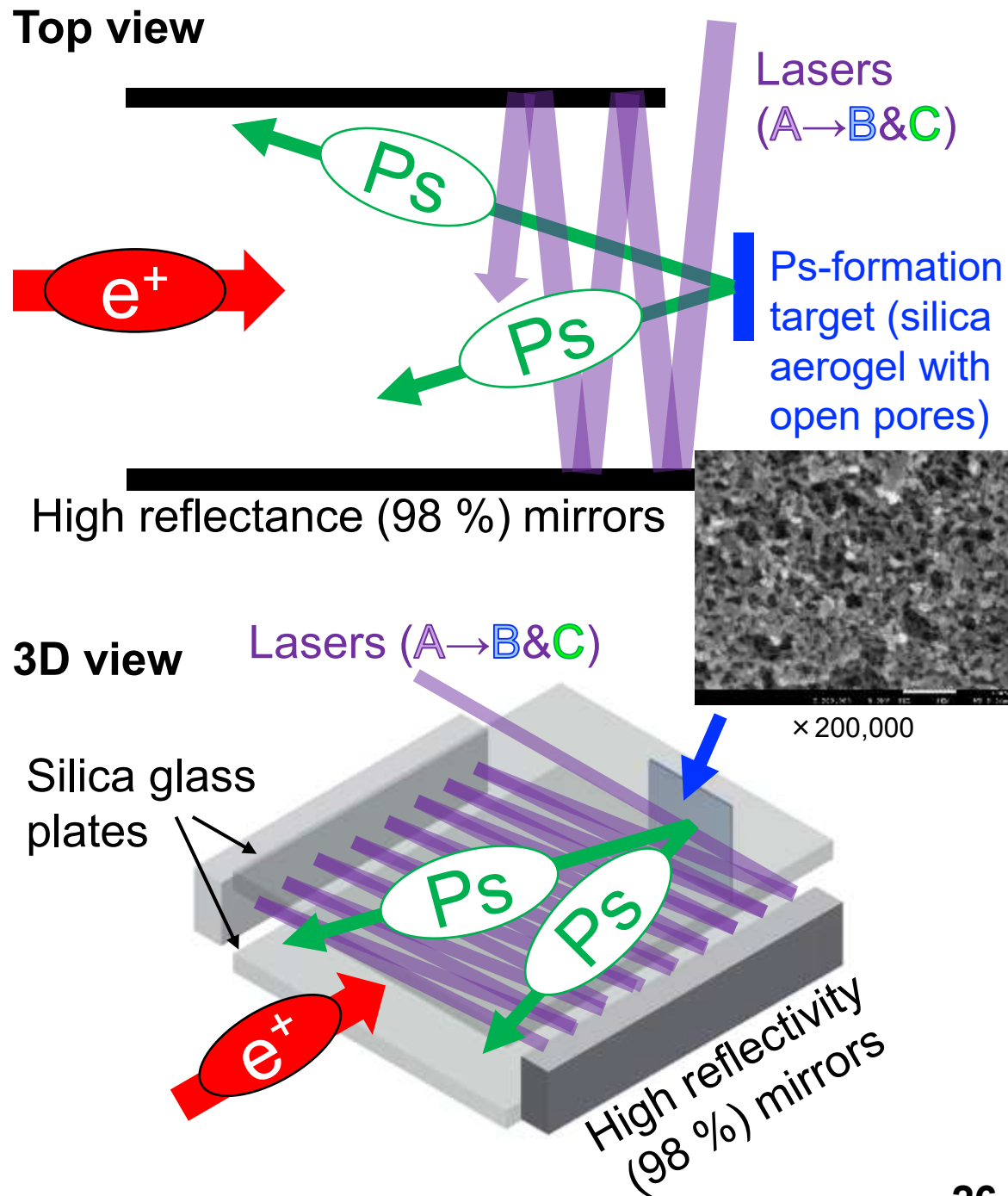


### (c)



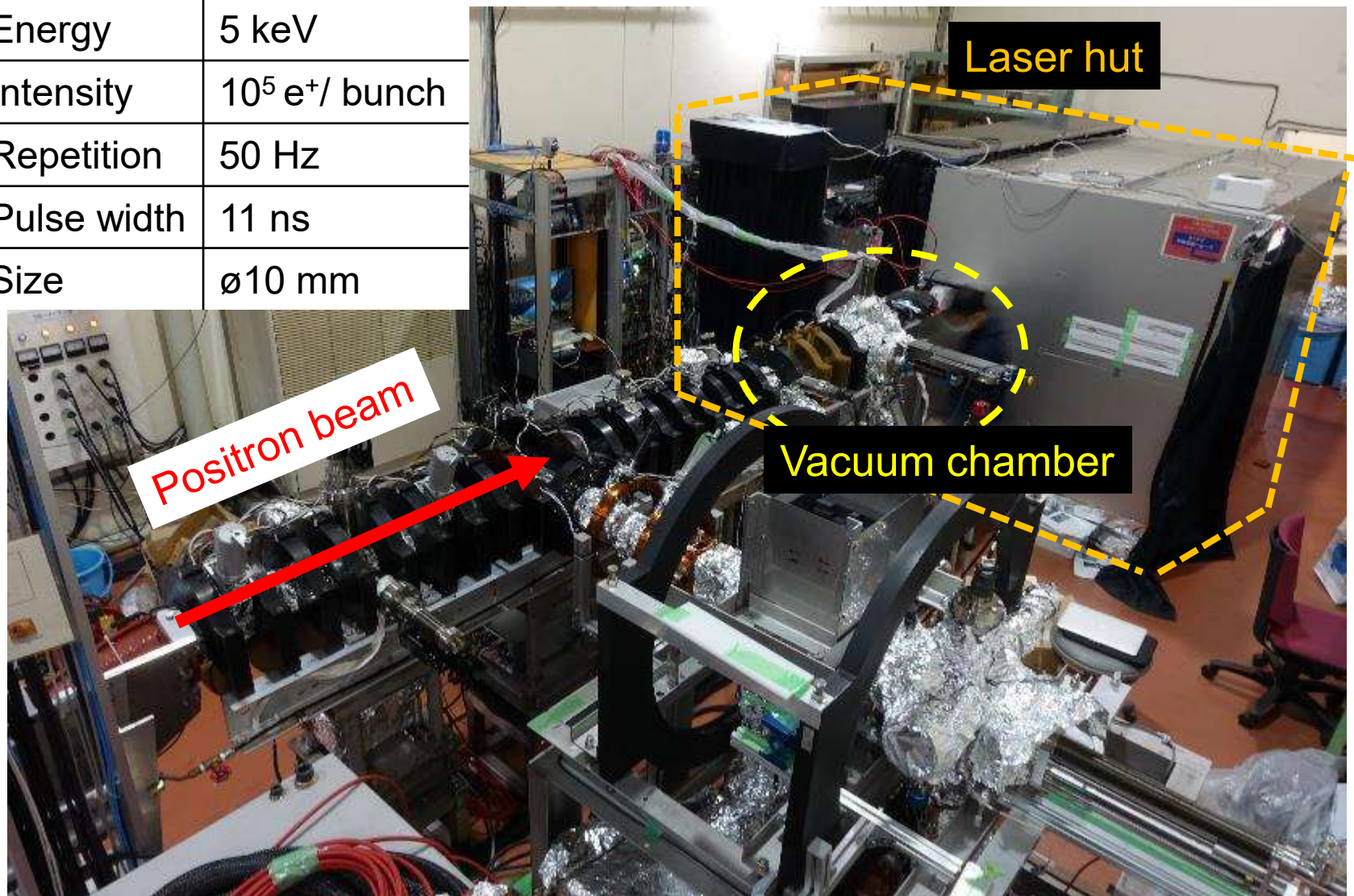
# 1D laser cooling of Ps in vacuum.

- Create Ps by irradiating a Ps-formation target (silica aerogel with open pores) with positron beam.
- Irradiate Ps emitted from the target with 243 nm UV laser.
- Reflect lasers for multiple times by high reflectivity mirrors to obtain the interaction area between the laser and Ps.
- Confine Ps with two silica glass plates.



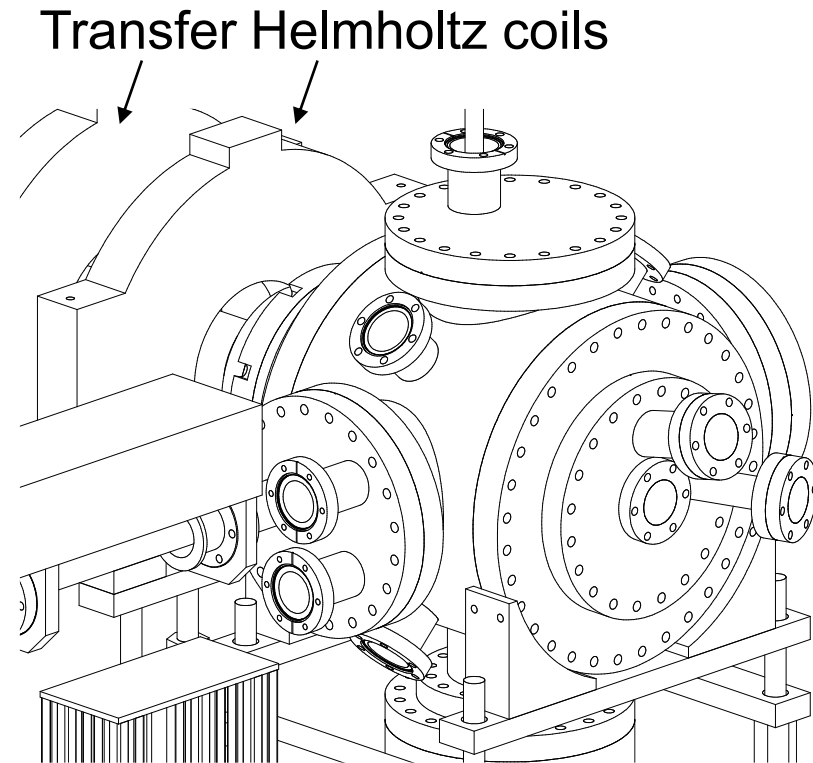
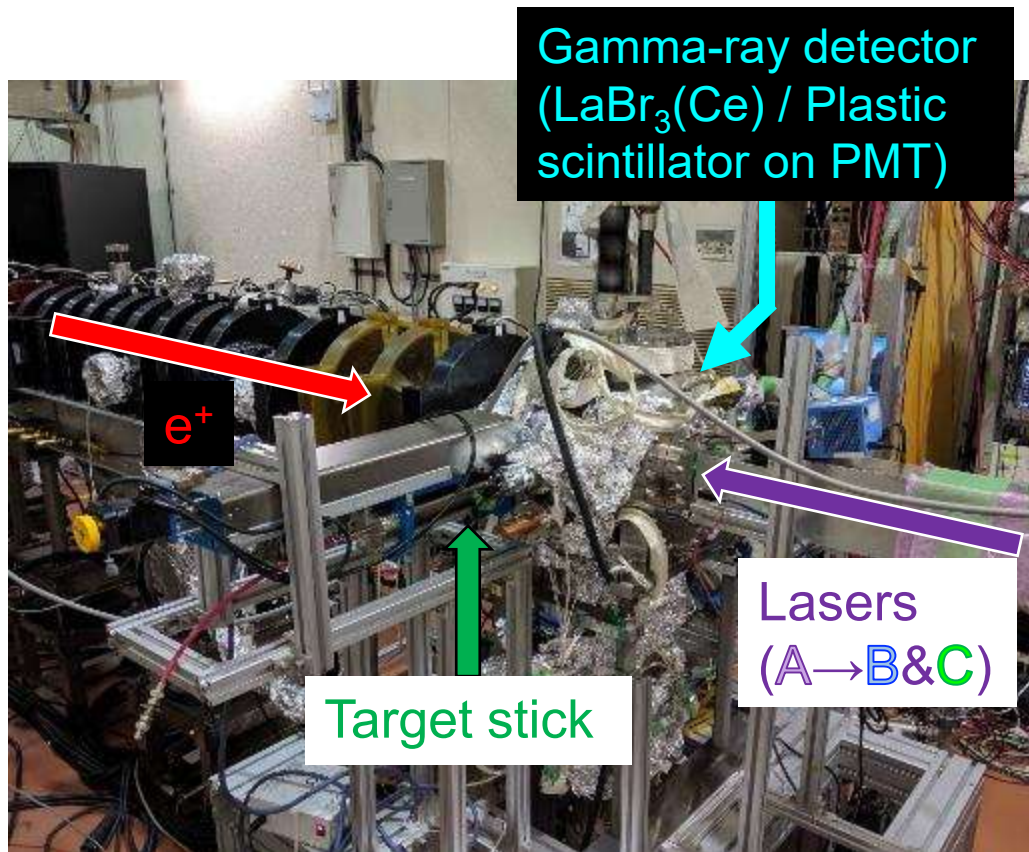
# Experimental setup at KEK-SPF

Energy	5 keV
Intensity	$10^5$ e <sup>+</sup> / bunch
Repetition	50 Hz
Pulse width	11 ns
Size	ø10 mm

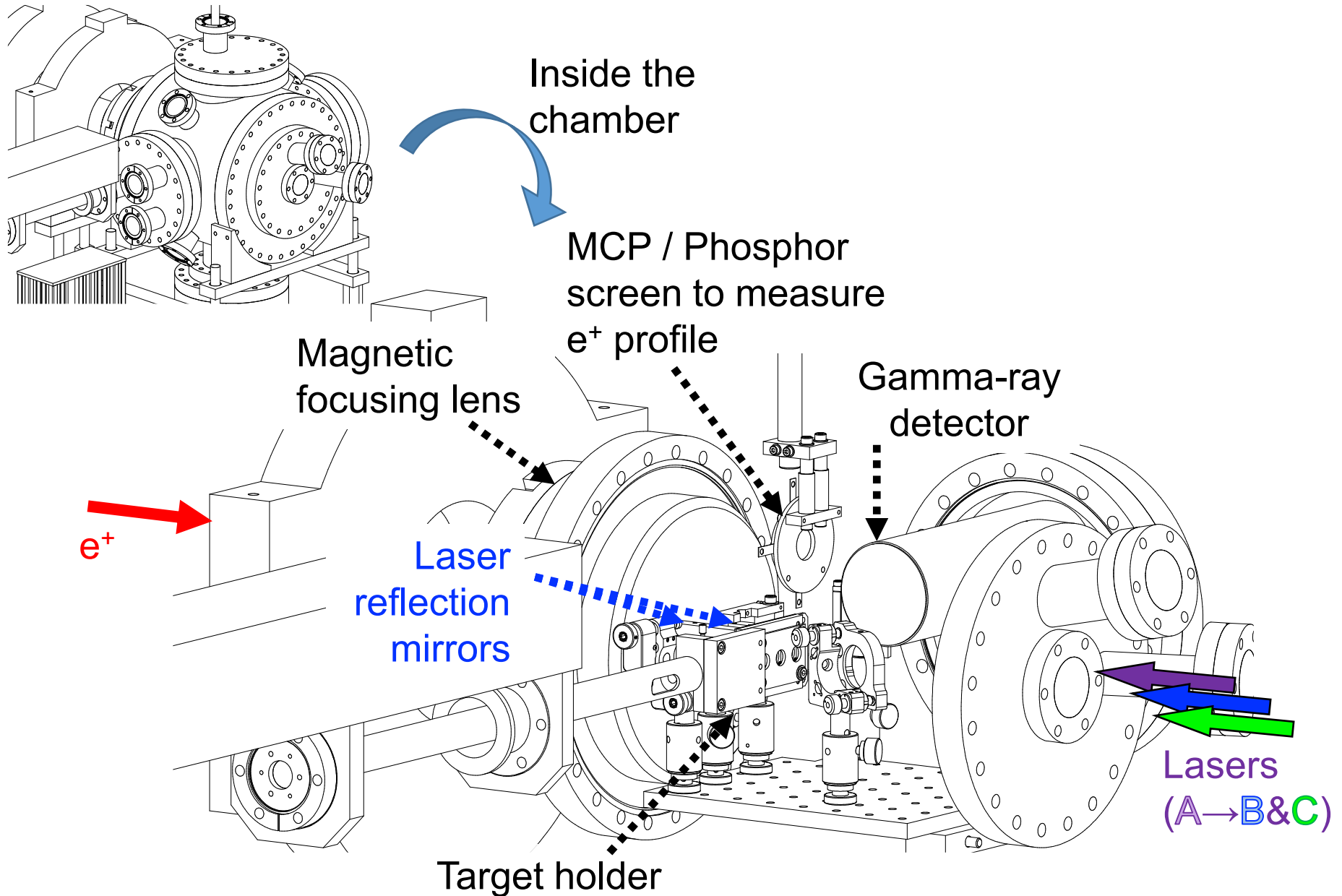




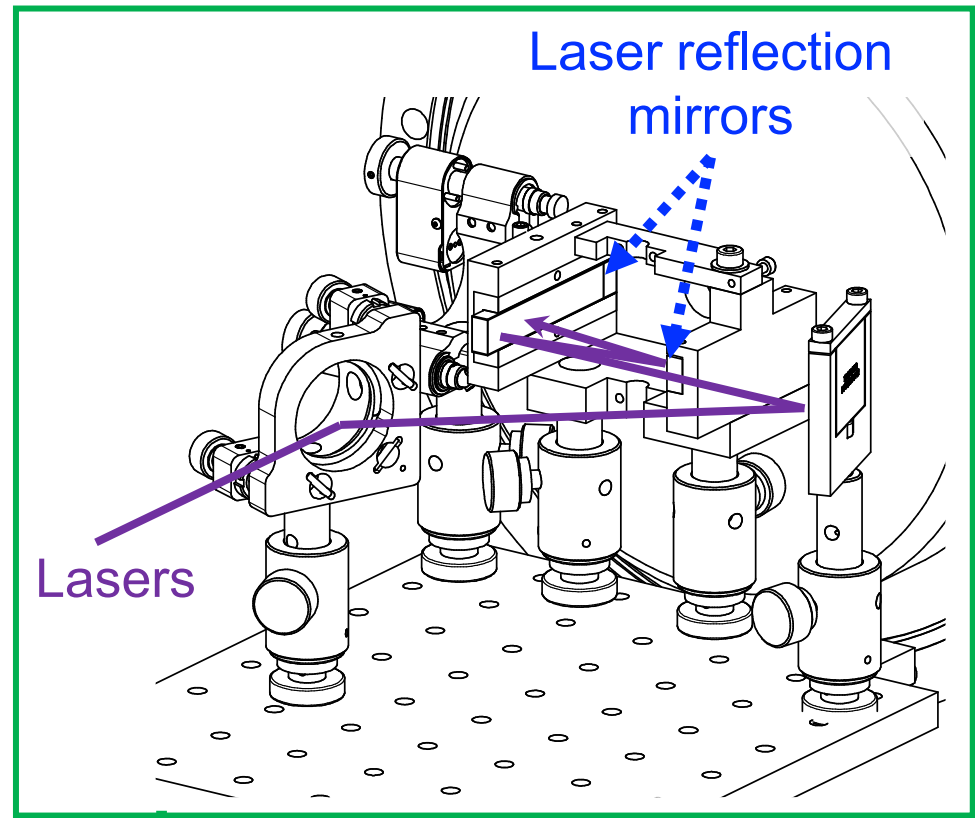
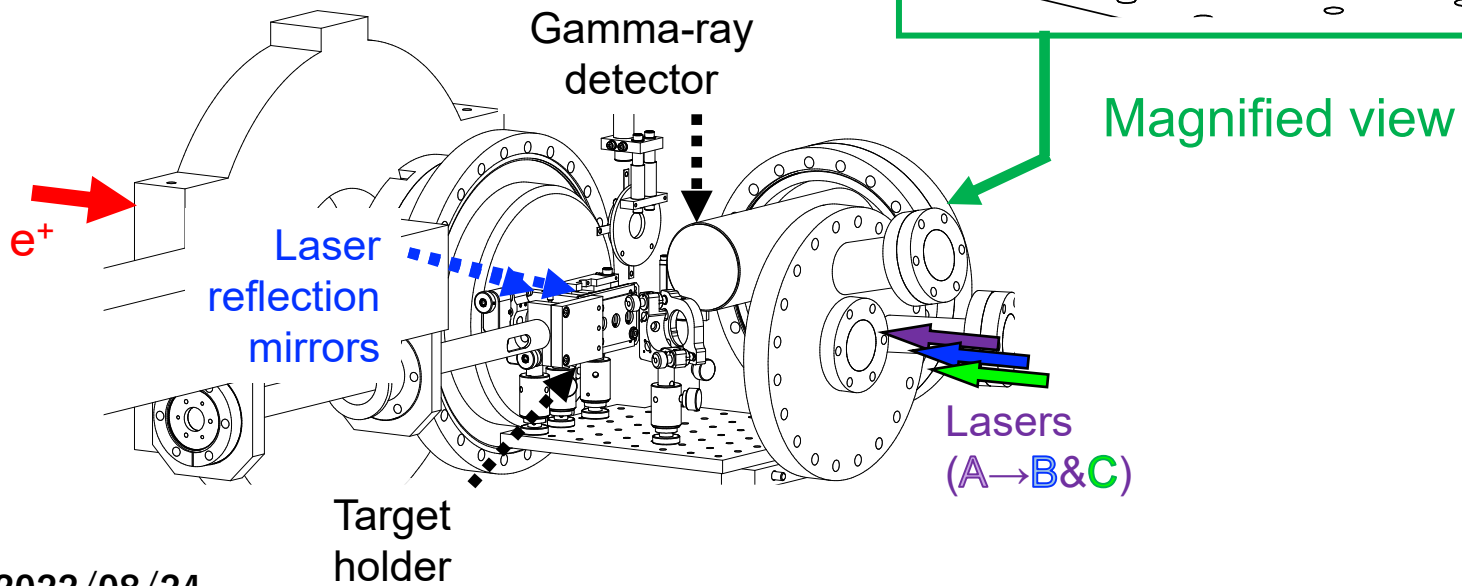
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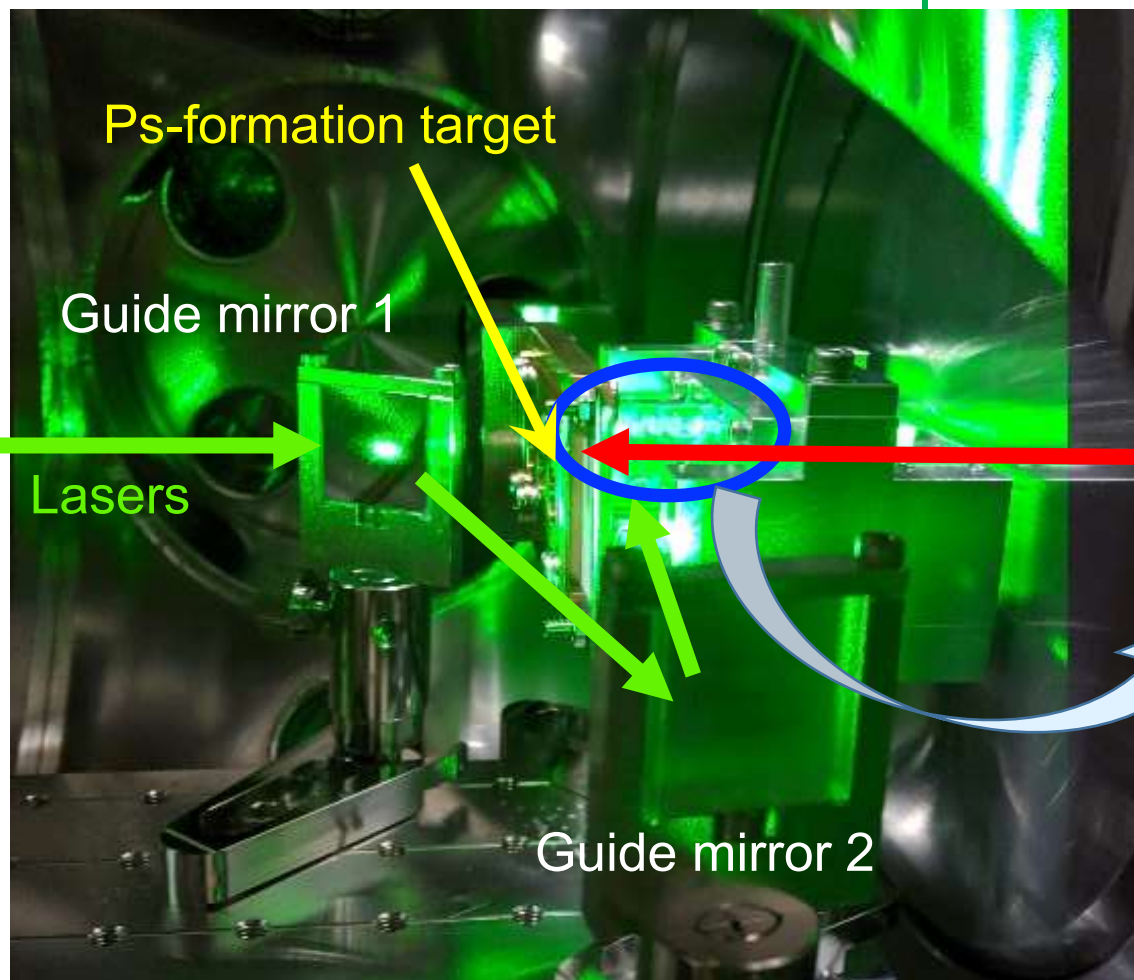
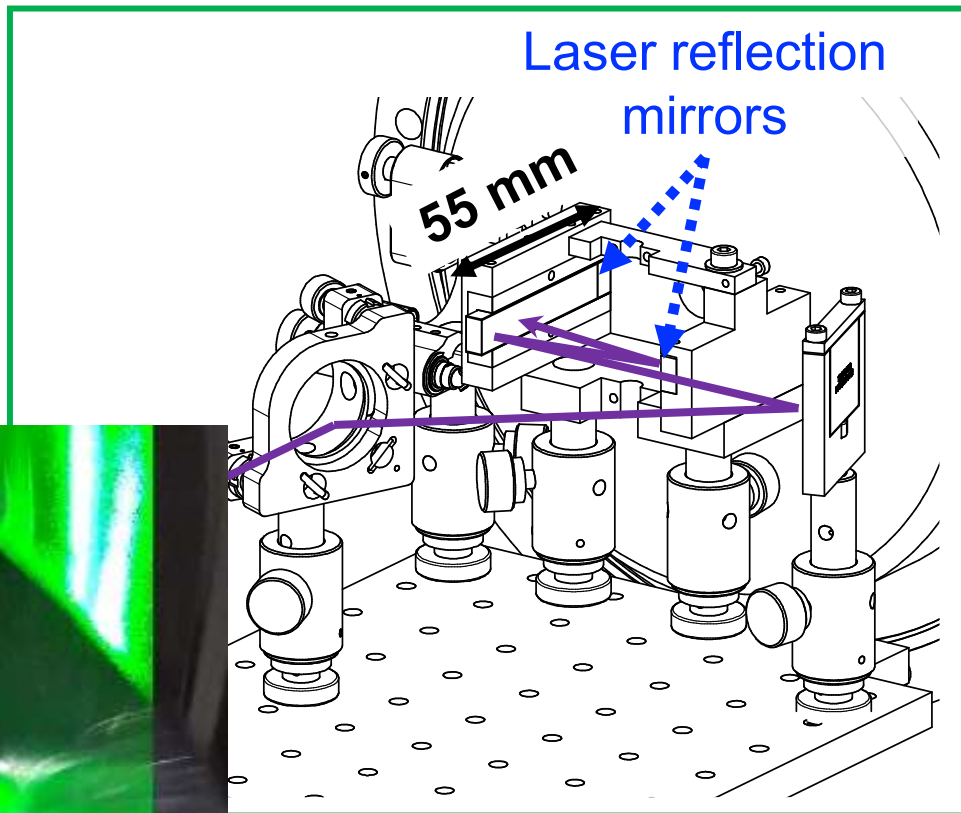
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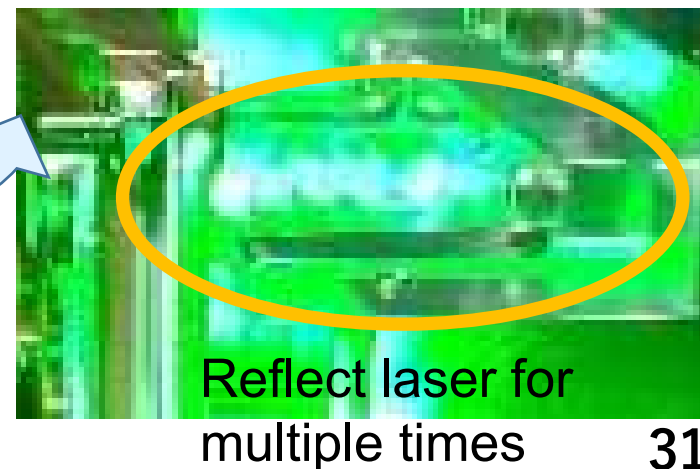
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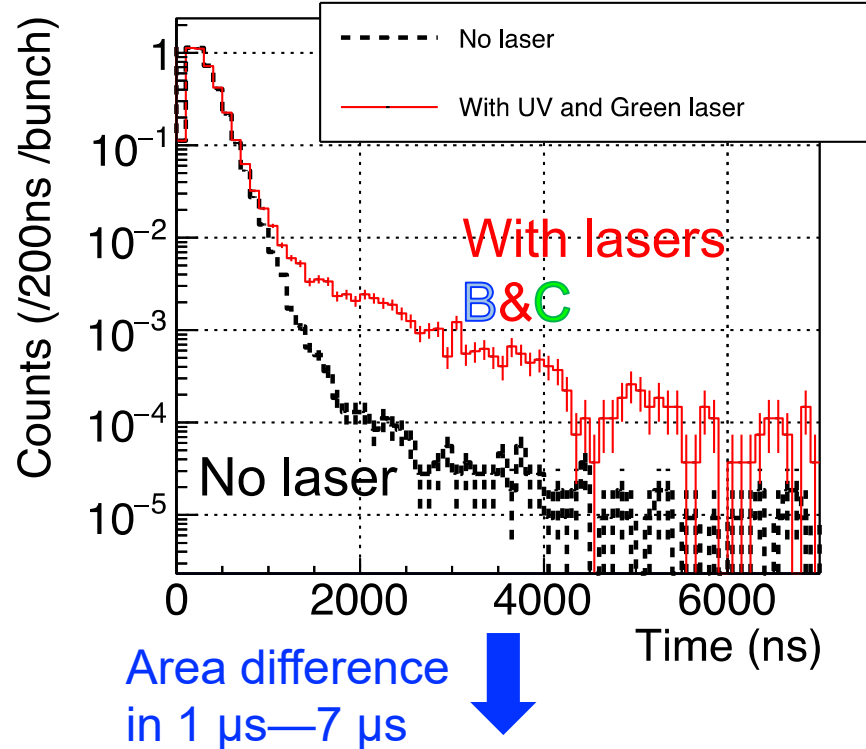
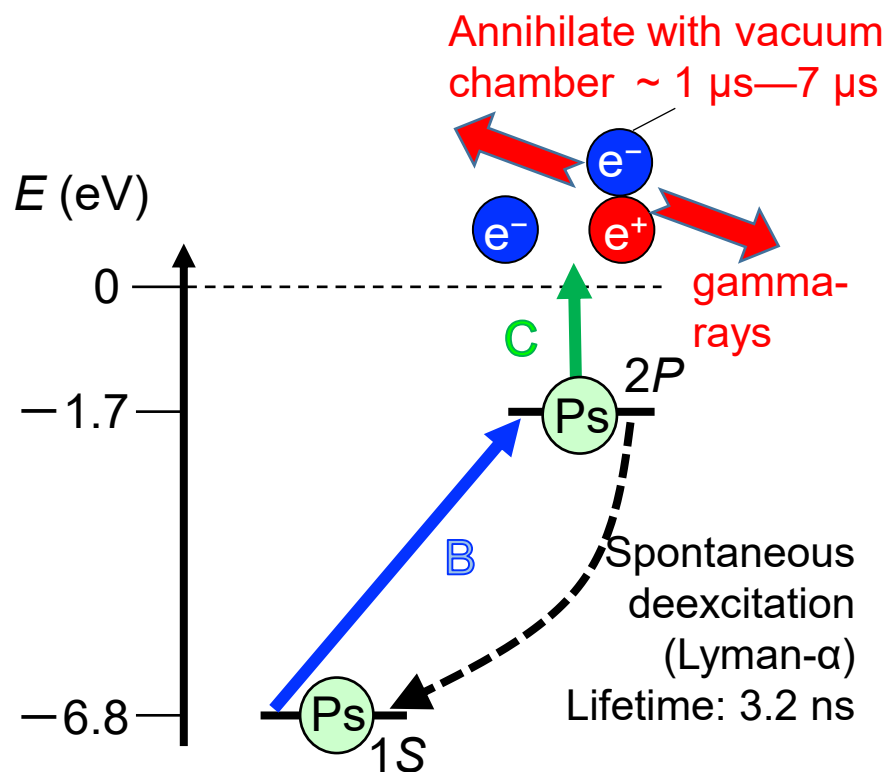
# Experimental setup at KEK-SPF



Positrons

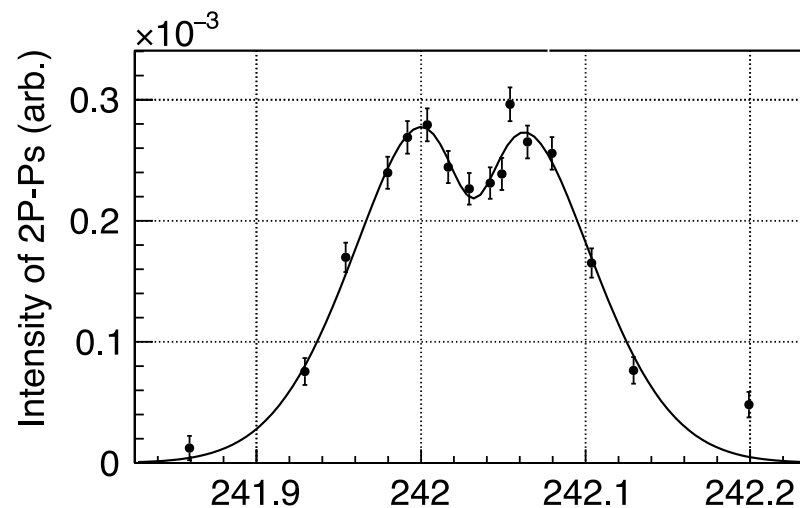
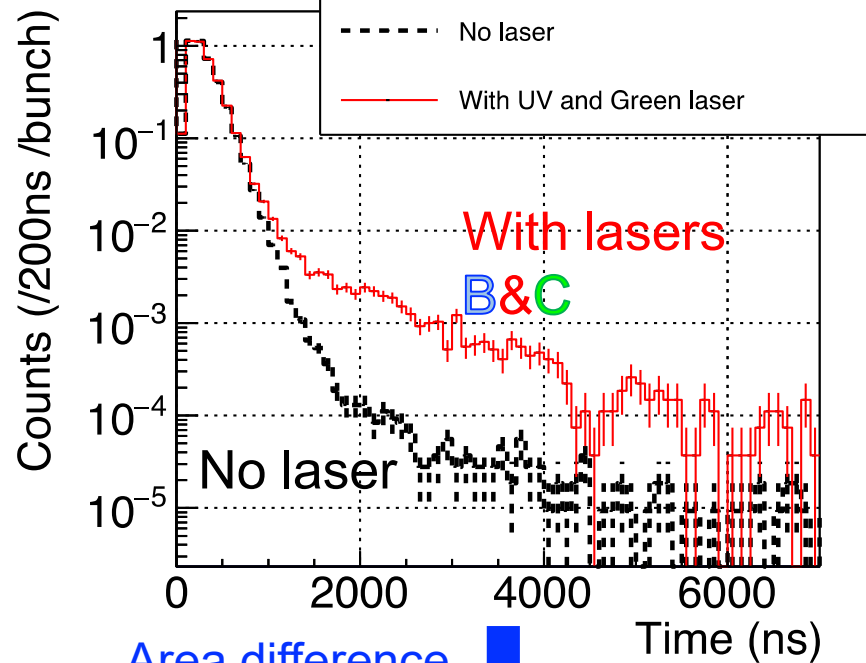
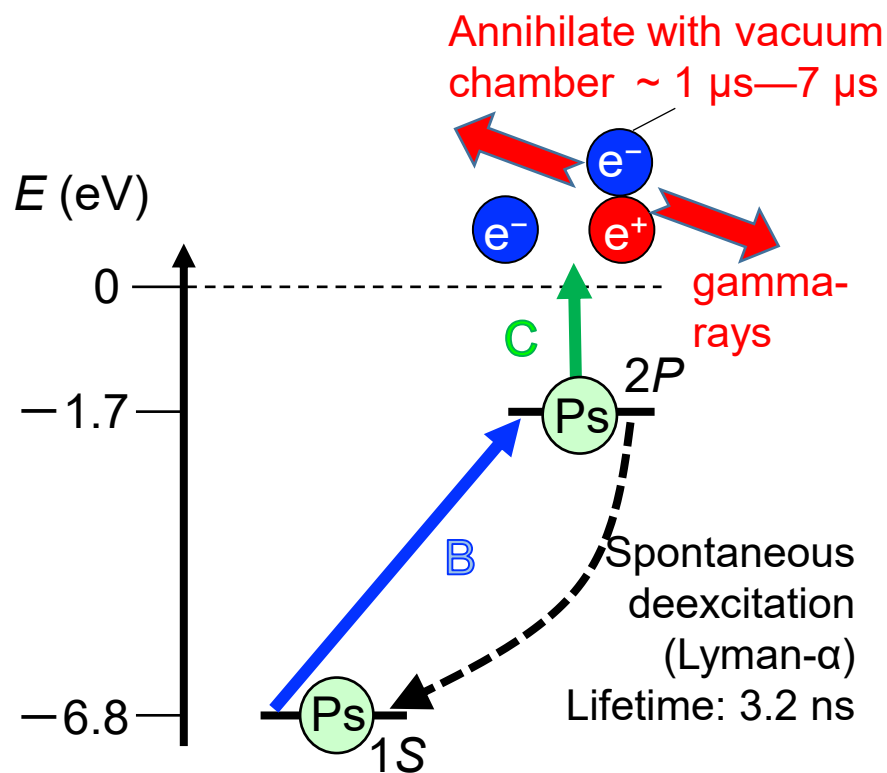


Without shining the cooling laser A, we confirmed stimulations of  $1S \rightarrow 2P$  transition of Ps with probe lasers B&C.





Without shining the cooling laser A, we confirmed stimulations of  $1S \rightarrow 2P$  transition of Ps with probe lasers B&C.

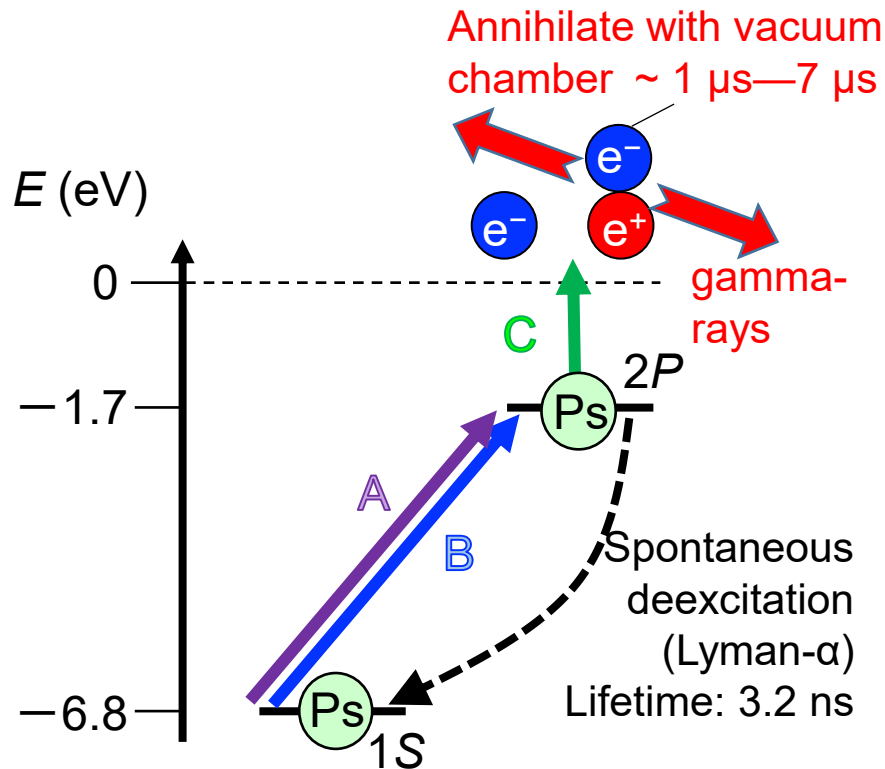


A. Ishida *et al.*, Photon Factory Activity Report 37(2020)201 (Japanese).

Wavelength of probe laser B (before calibration, nm)

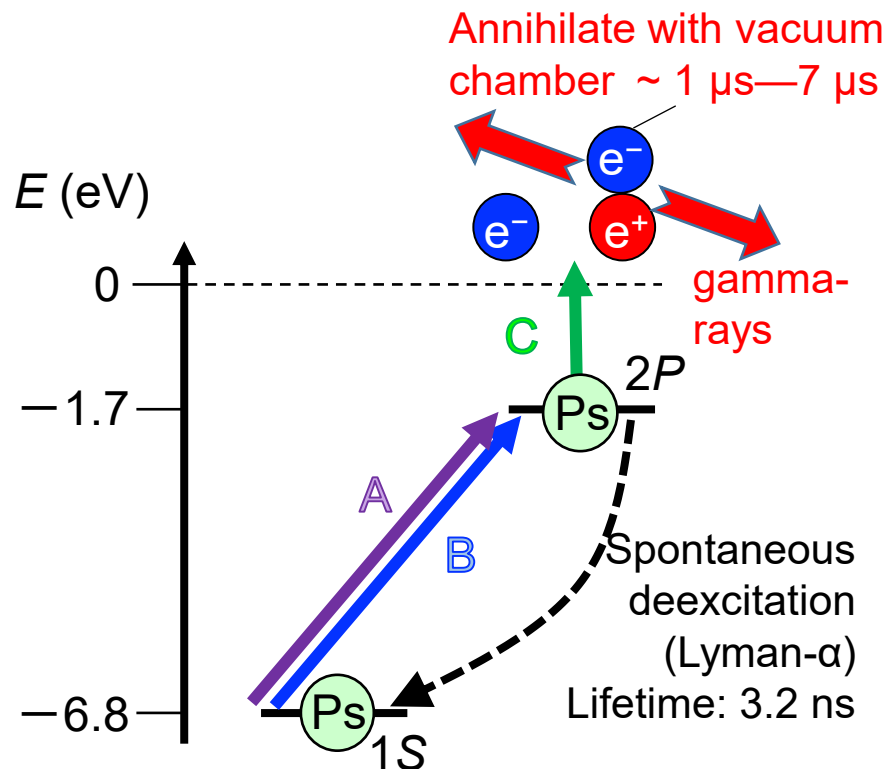
The Ps  $1S \rightarrow 2P$  transition signal was nearly **doubled at all wavelengths of probe laser B** when we did the **first trial to shine a prototyped Ps cooling laser A** before shining lasers **B&C**.

→ We still need further detailed analysis and consolidations, but so far it turned out that there was a **huge systematic effect caused by shining UV lasers to the apparatus**, especially to the laser reflection mirrors.



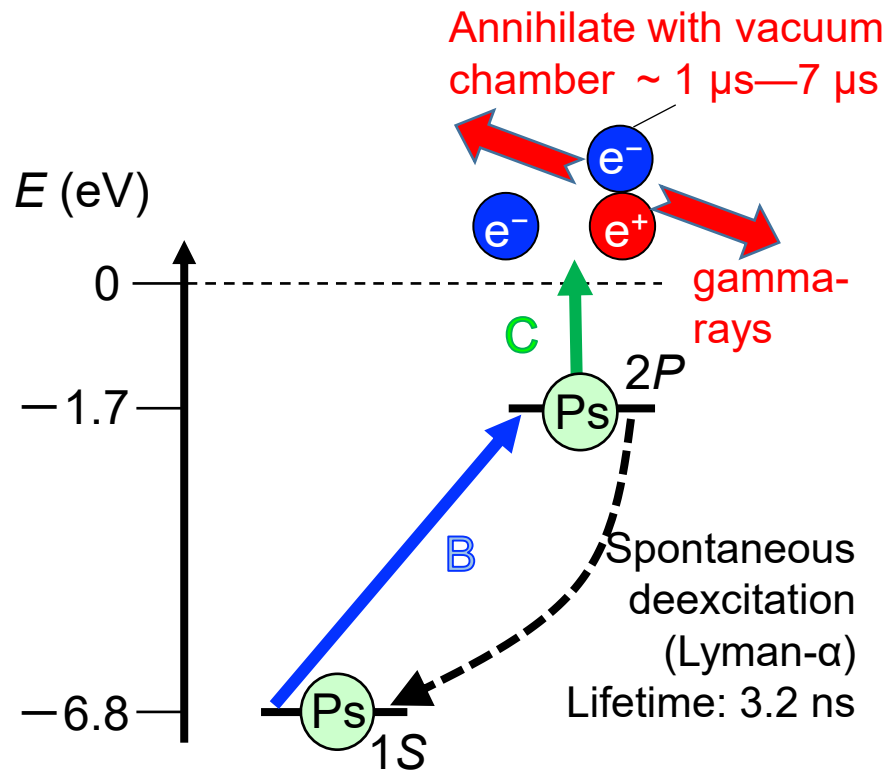
To investigate possible systematic effects caused by shining UV lasers to the apparatus, **we removed the cooling laser A**.

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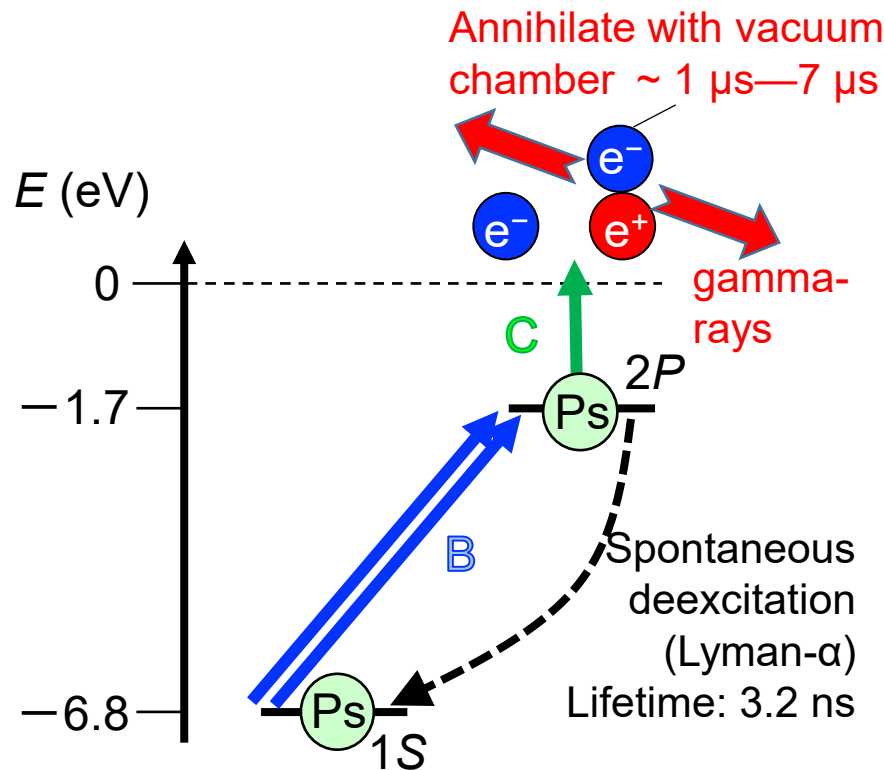
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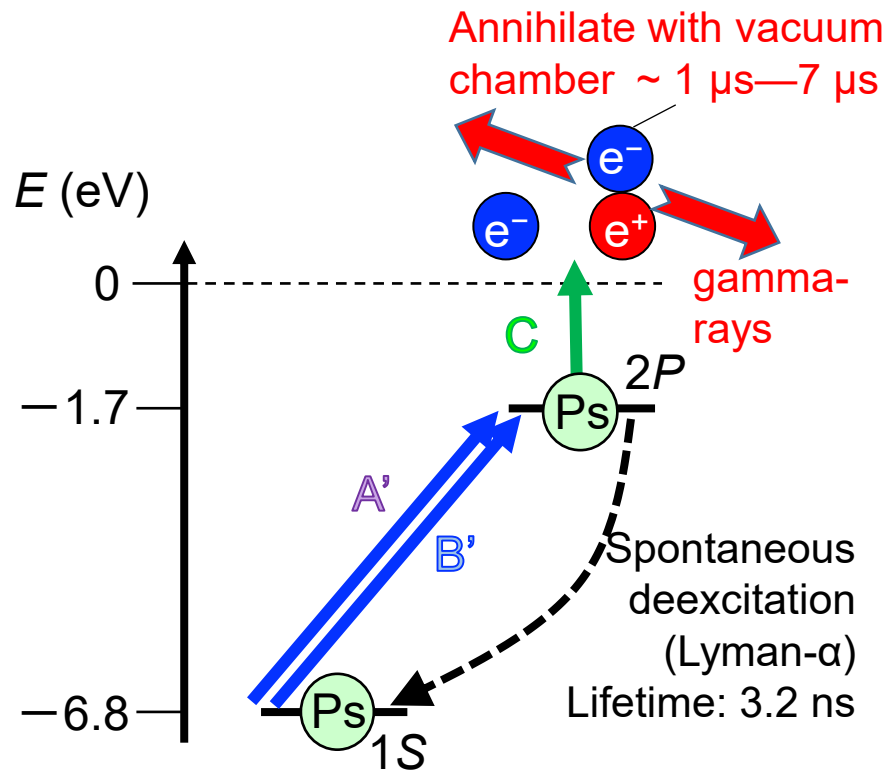
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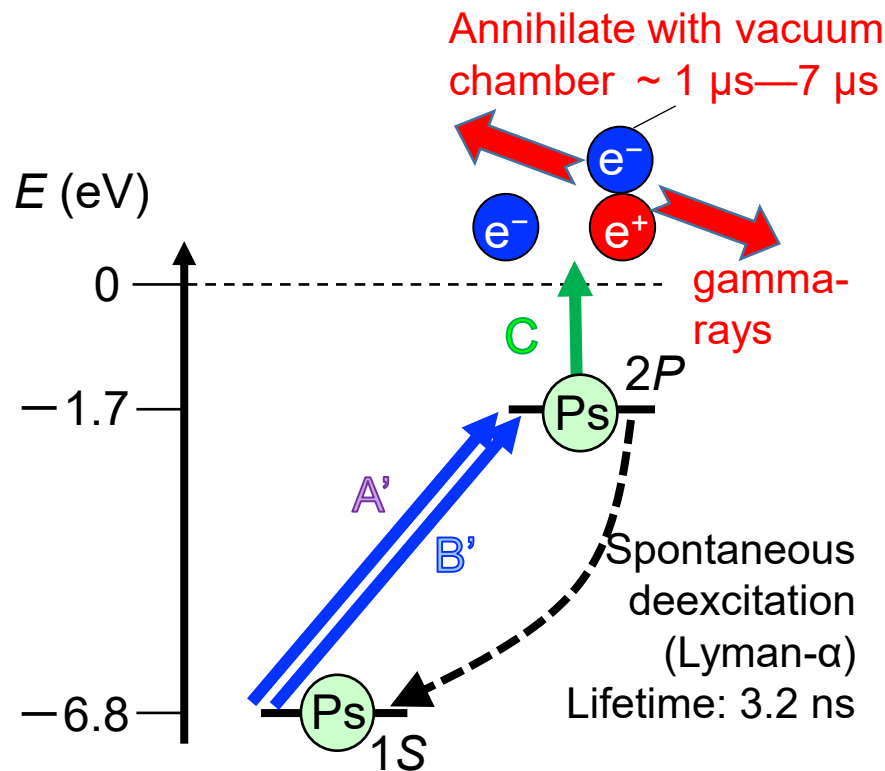
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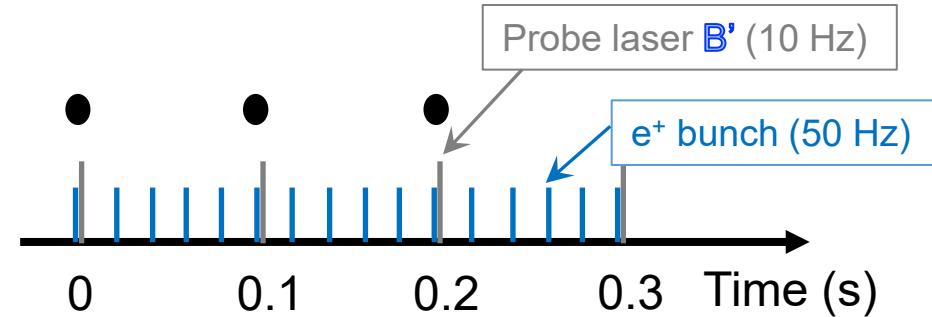


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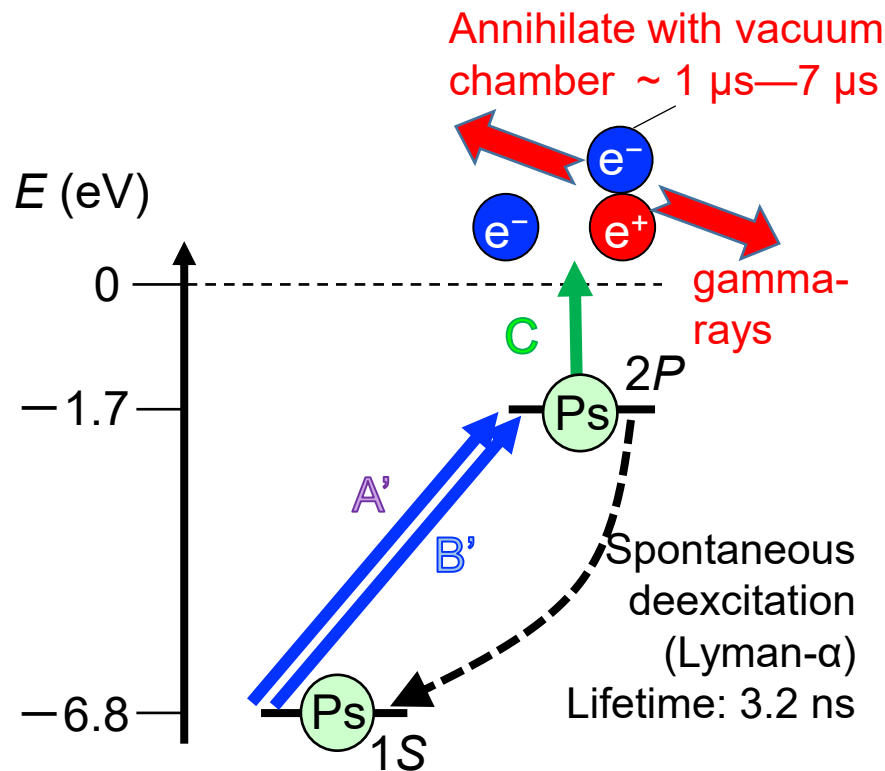


(a) Without *mock* laser **A'**

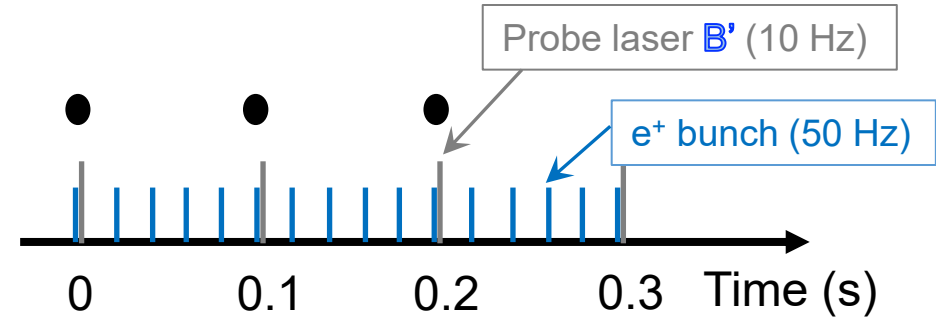


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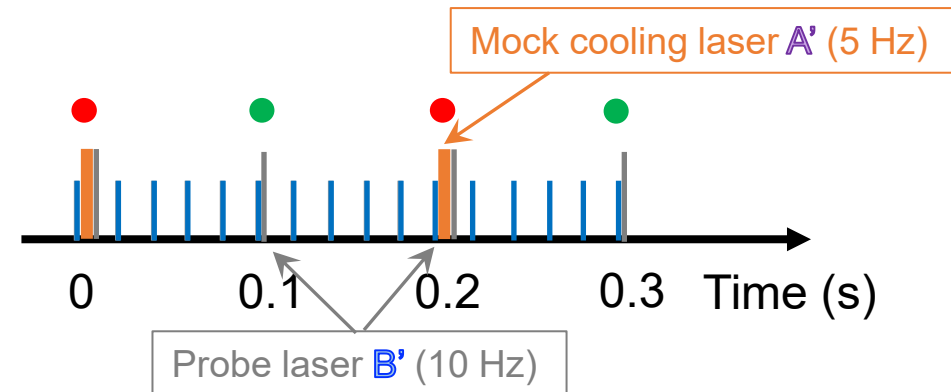
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(a) Without *mock* laser **A'**



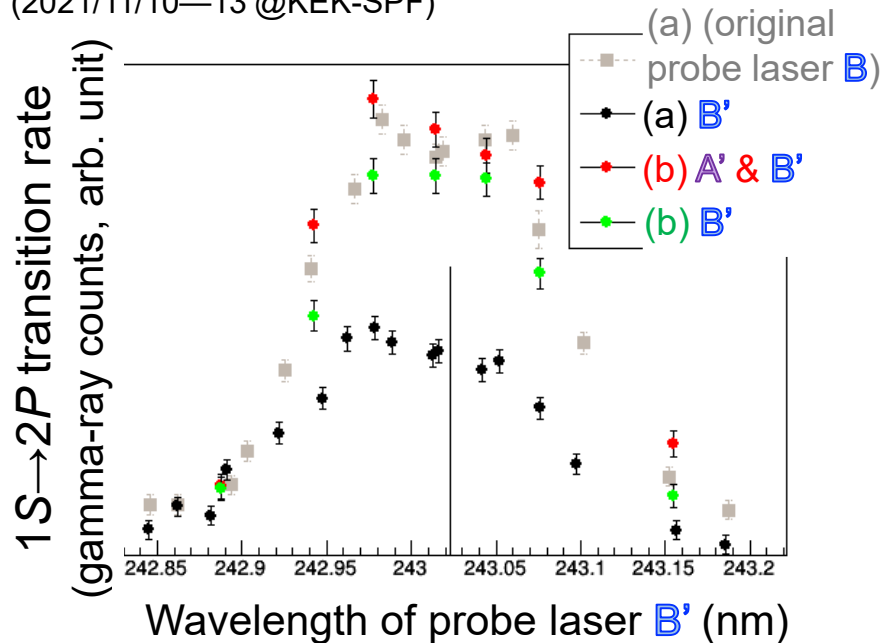
(b) With shining *mock* laser **A'** at 5 Hz.





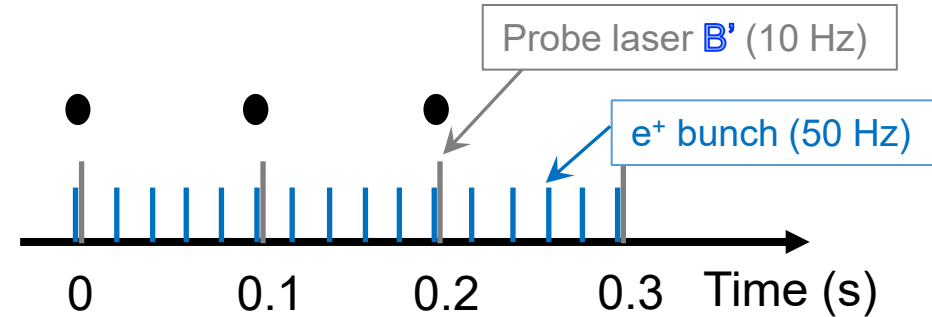
We observed a **signal increase by the *mock* cooling experiment**.  
 The systematic effect by shining the *mock* laser A' could be seen  
**even at the timings when the mock laser A' was not shined**.  
 The decay constant of the effect was found to be  $\sim 5$  minutes.

(2021/11/10—13 @KEK-SPF)

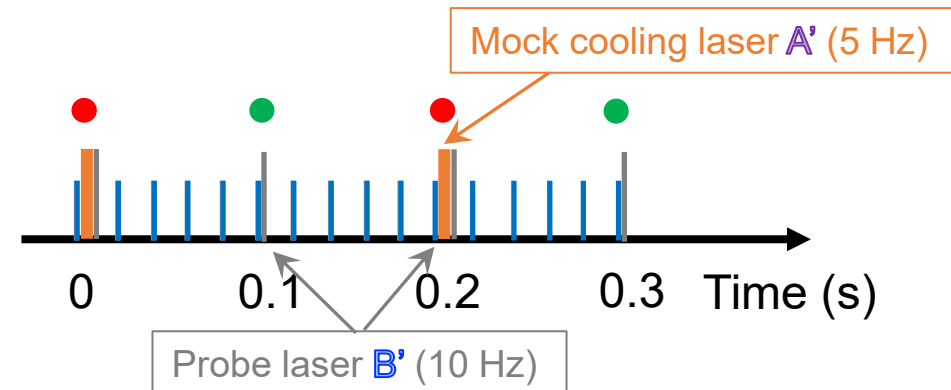


The transition rates are high for both conditions in experiment (b).

(a) Without *mock* laser A'



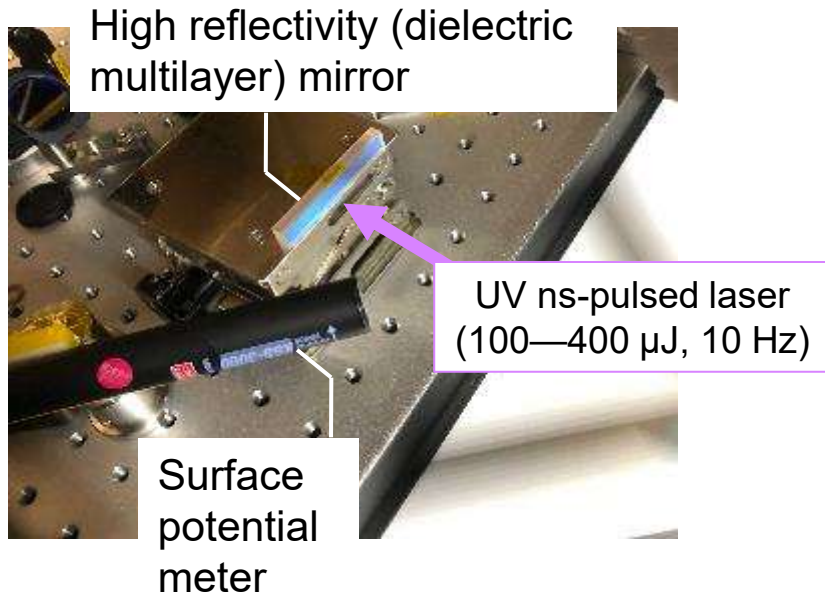
(b) With shining *mock* laser A' at 5 Hz.



A. Ishida *et al.*, Photon Factory Activity Report **39**(2022)183 (Japanese).

We found that **the laser reflection mirrors were negatively charged by UV lasers**, which caused a huge systematic bias on the measurement of Ps transition rate.

Electric potential measurement  
in atmosphere.

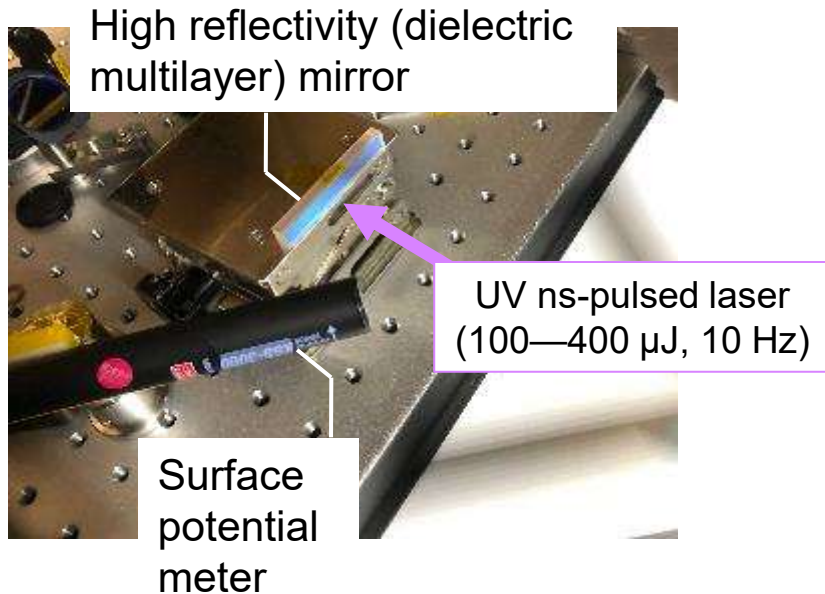


The mirror was negatively charged.

→ The trajectories of ionized positrons were changed by the electric field, which caused a huge bias on the detection efficiency of Ps transitions.

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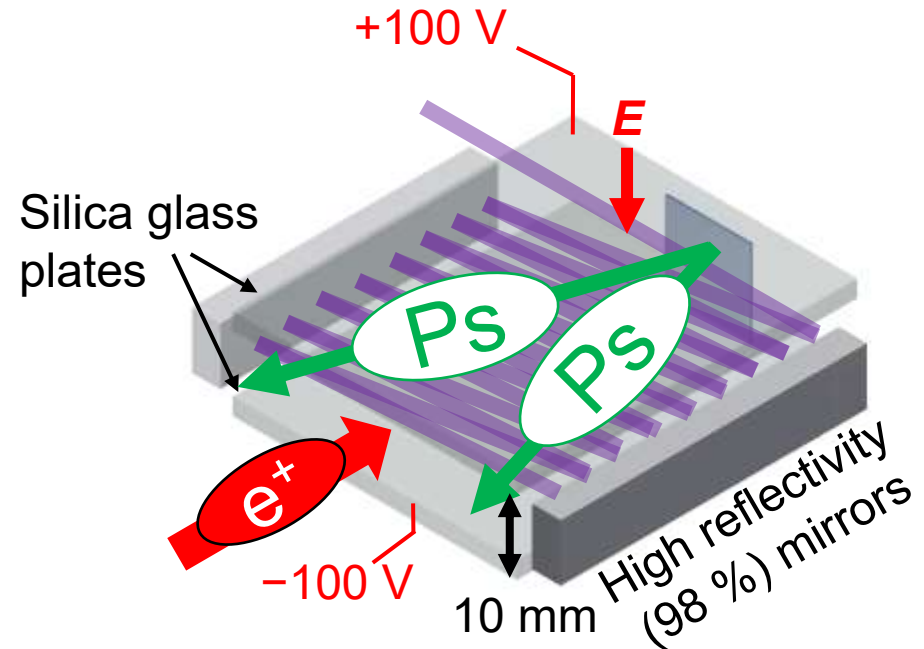
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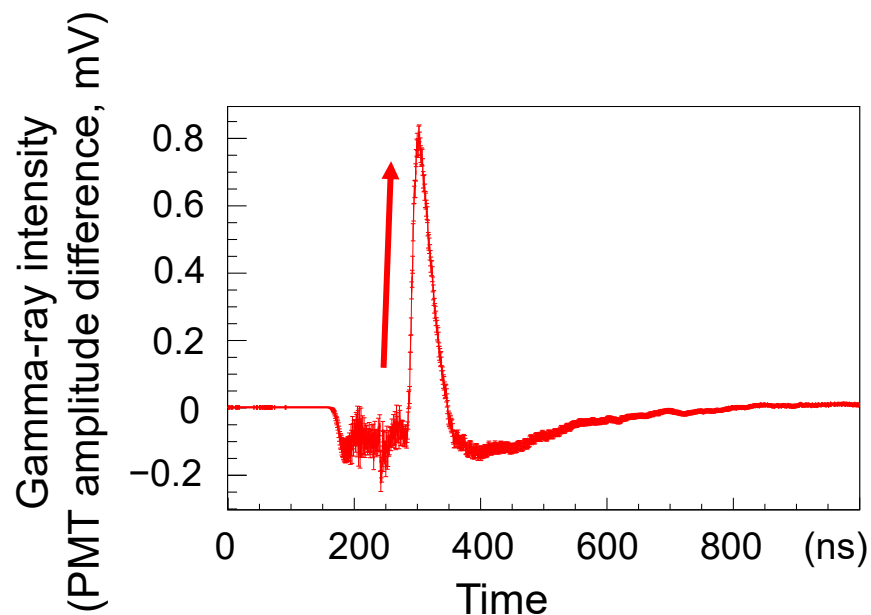
→ The trajectories of ionized positrons were changed by the electric field, which caused a huge bias on the detection efficiency of Ps transitions.

Our solution:



We attached tungsten (W) meshes (93 % open area) to form an electric field. → Control the trajectories of ionized positrons.

We observed a **sharp increase of the gamma-ray counts** at the timing of laser irradiation.

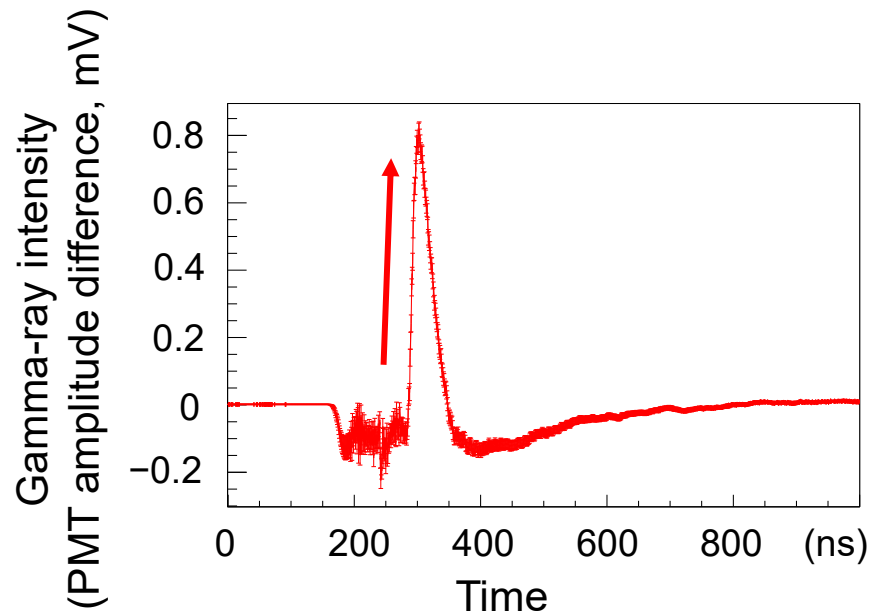


Forced positrons to annihilate immediately after Ps ionization so that we obtain  $1S \rightarrow 2P$  transition rate with a timing window which is as narrow as a few ns.

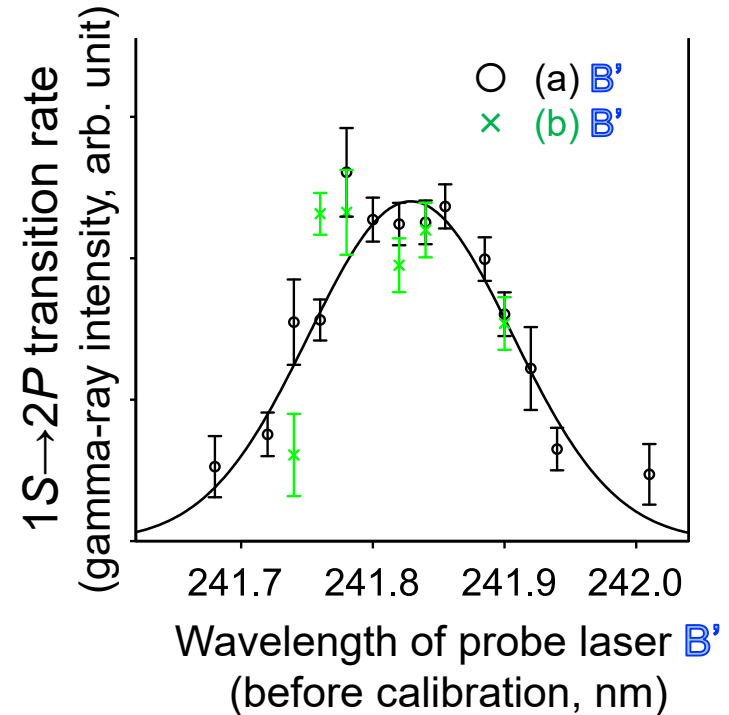
We observed a **sharp increase of the gamma-ray counts** at the timing of laser irradiation.

We **succeeded in removing the systematic effect** caused by the mirror charged by the UV lasers.

We expect a proof-of-principle experiment of Ps laser cooling within a year.



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Checked by another mock cooling experiment.  
No difference was observed.

# Summary

We want to cool Ps down to 10 K for various physics interests.

1. Combination of Ps thermalization cooling and laser cooling will be efficient enough to realize Ps-BEC.
2. We have developed a dedicated laser system for Ps cooling, Chirped Pulse-Train Generator (CPTG), with 300-ns pulse duration and  $0.4 \text{ GHz ns}^{-1}$  frequency chirp.
3. We are trying a proof-of-principle experiment to laser-cool Ps in vacuum at KEK-SPF. We succeeded in removing the systematic effect caused by the mirror charged by the UV lasers.
4. We expect a proof-of-principle experiment of Ps laser cooling within a year.

K. Shu, X. Fan, T. Yamazaki *et al.*, J. Phys. B: At. Mol. Opt. Phys., **49**, 104001 (2016).

A. Ishida, K. Shu, T. Murayoshi *et al.*, JJAP Conf. Proc., **7**, 011001 (2018).

K. Shu, A. Ishida, T. Namba *et al.*, Phys. Rev. A, **104**, L050801 (2021).

K. Yamada, Y. Tajima, T. Murayoshi *et al.*, Phys. Rev. Applied, **16**, 014009 (2021).

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