

# **Positronium Laser Cooling: Improving detector pulse analysis efficiency via machine learning**

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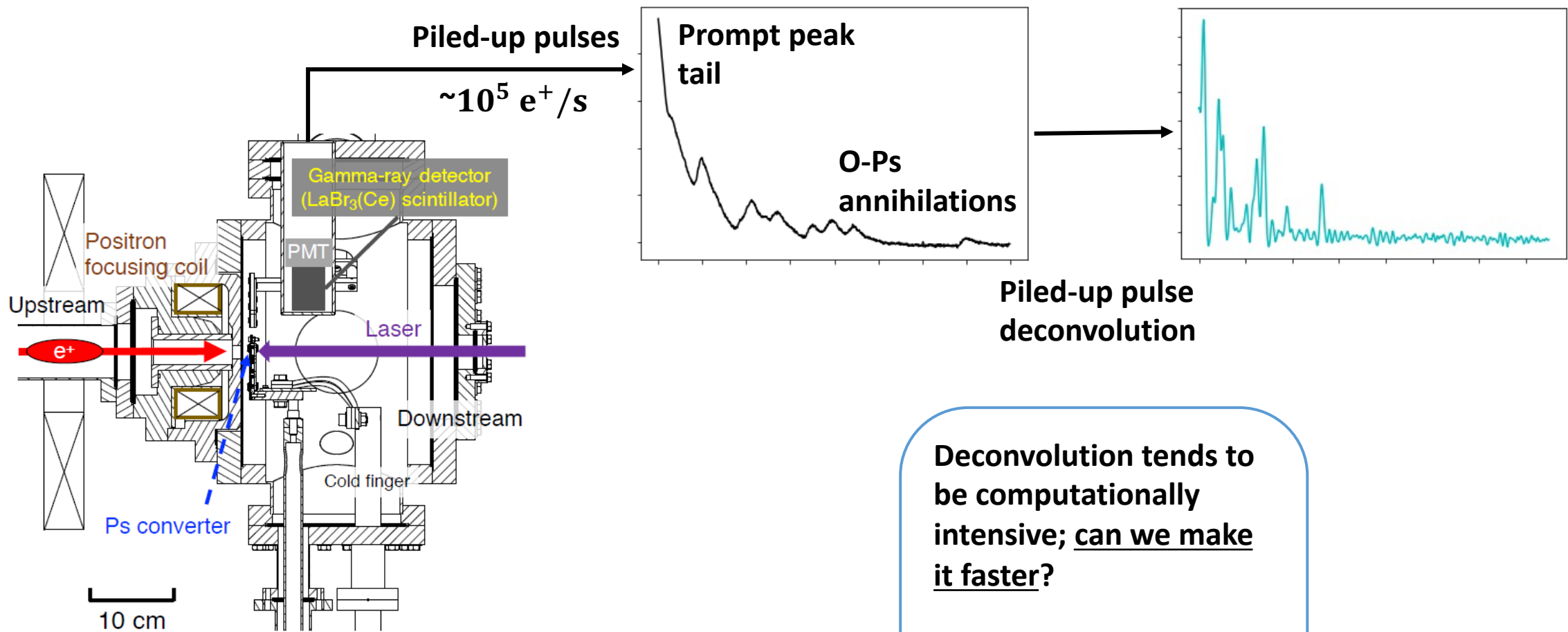
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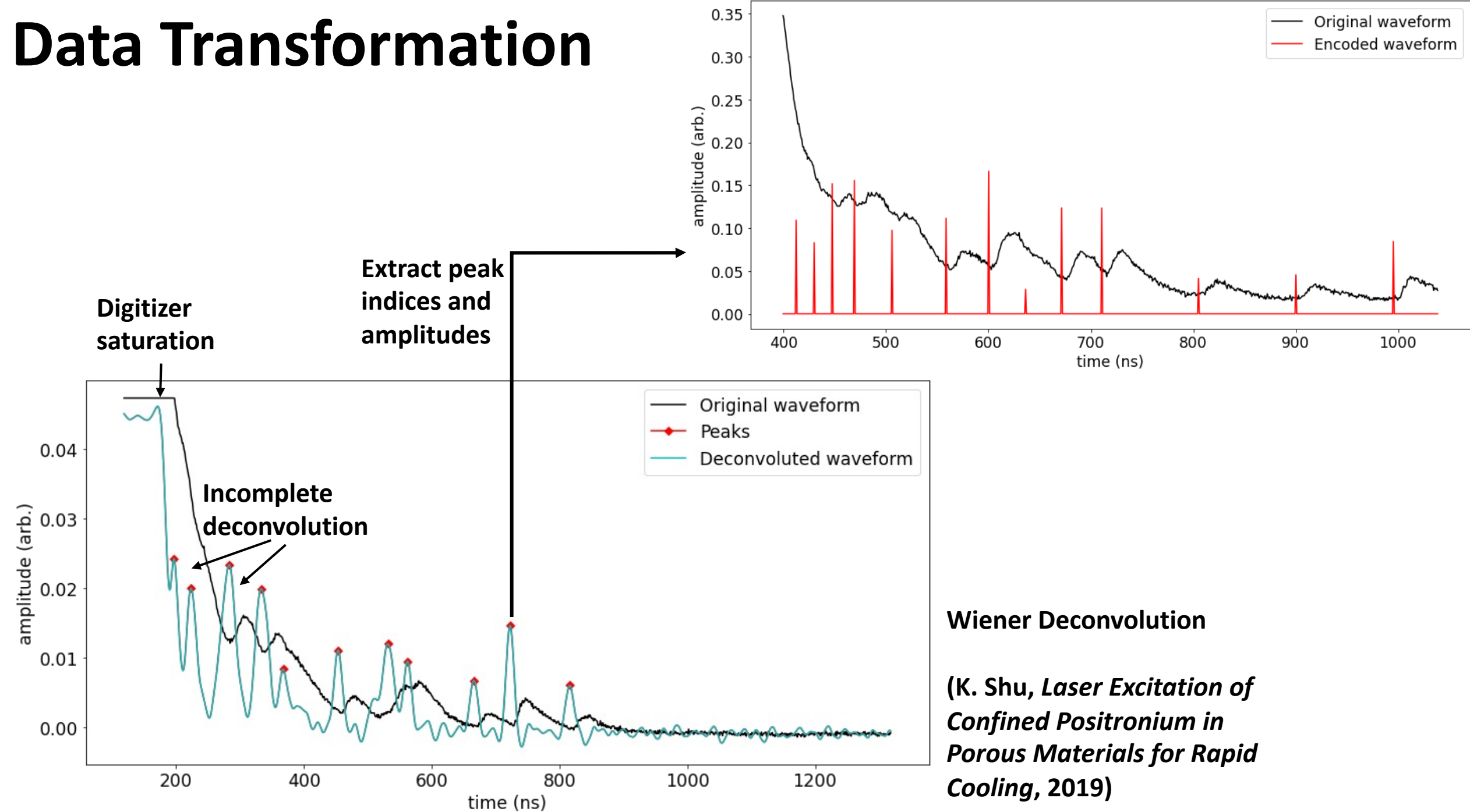
# Experiment and Motivation



Deconvolution tends to be computationally intensive; can we make it faster?

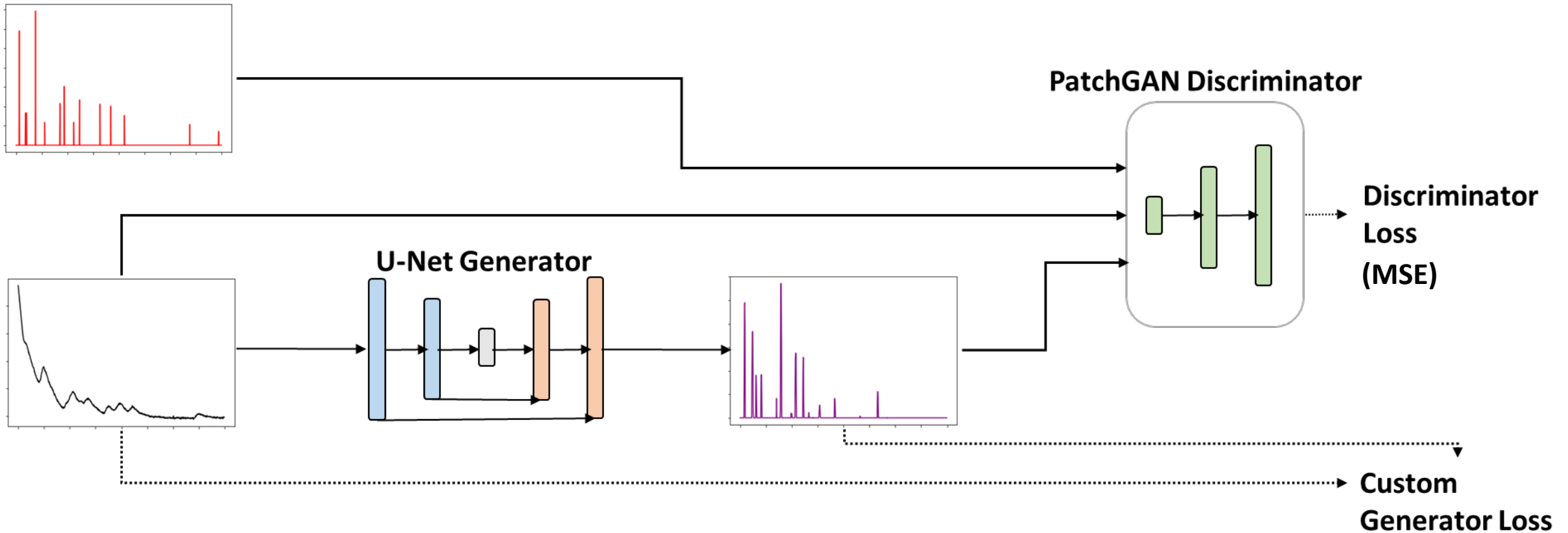
Machine learning inference time is generally fast compared to other methods

# Data Transformation



# Architecture and Training

Generator loss using only the difference (residuals or MSE) between waveforms is difficult to train when using delta-like peaks. Current loss:  $\sum |x - y| + c * \sum |x_{peaks} - y_{peaks}|$ . The loss function is almost (or just) as important as the architecture.



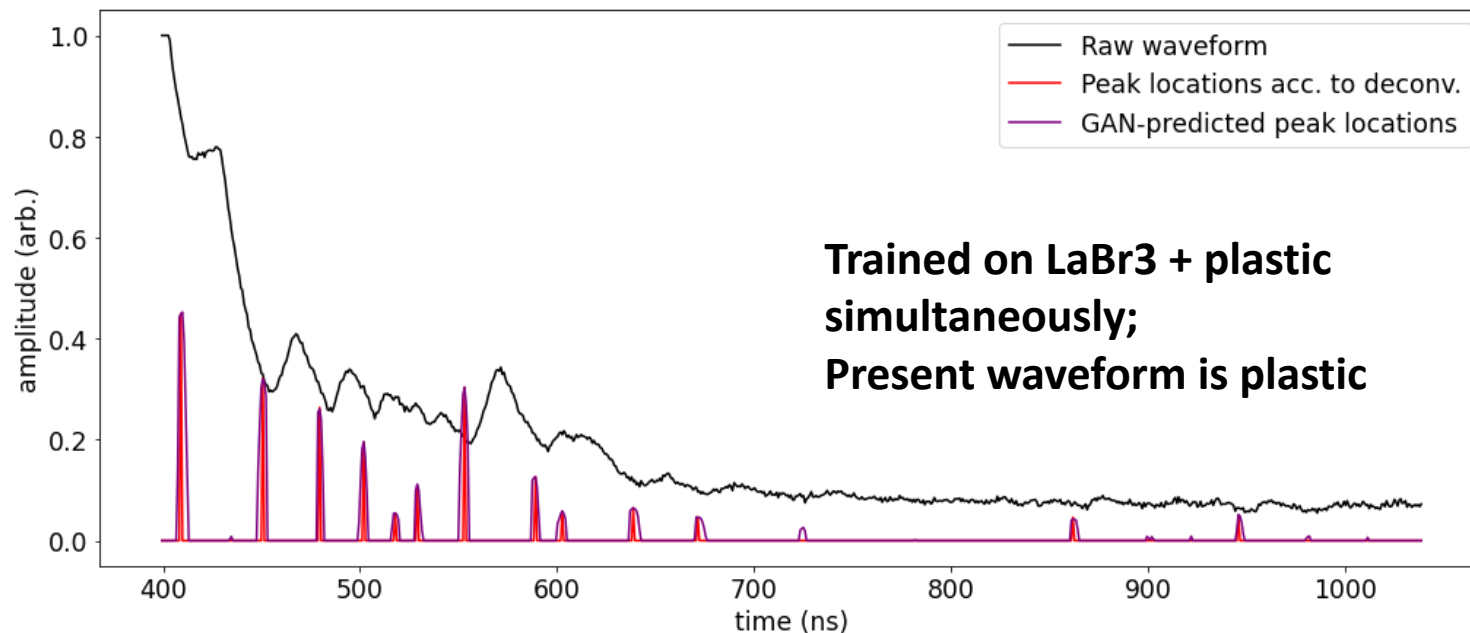
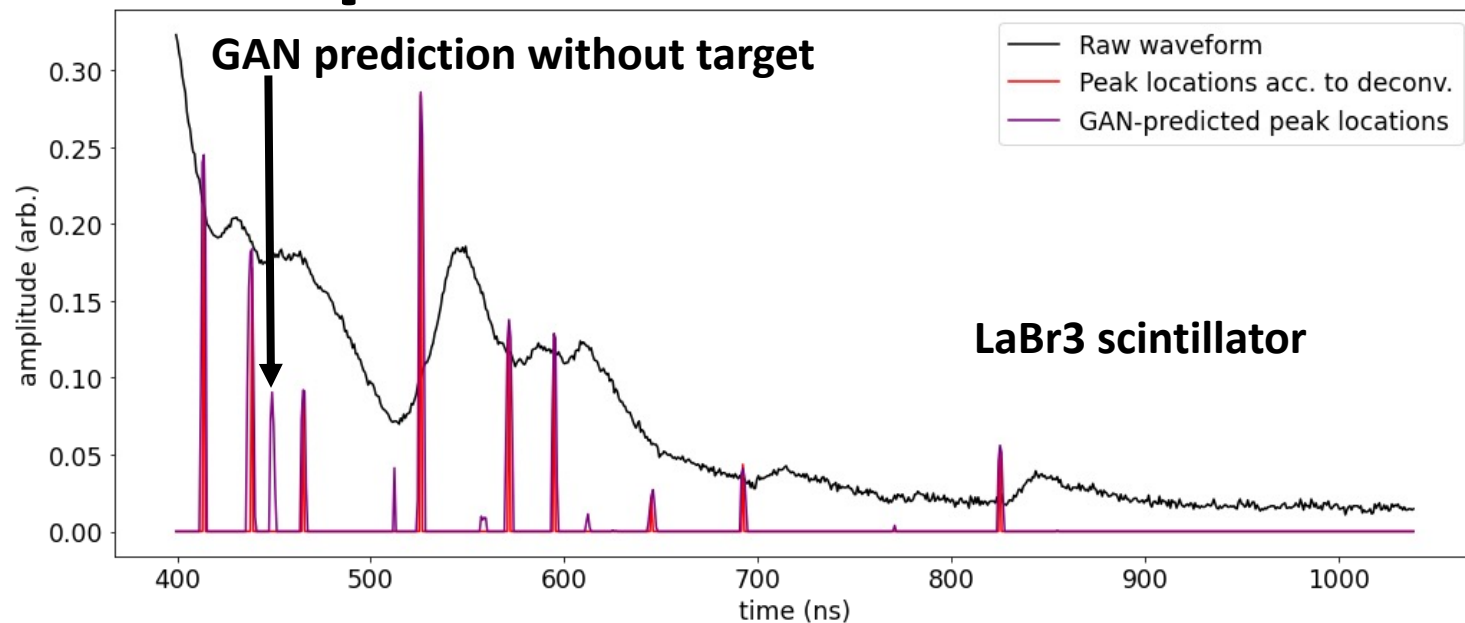
K. Schawinski, et al., *Generative adversarial networks recover features in astrophysical images of galaxies beyond the deconvolution limit*, 2017

<https://doi.org/10.1093/mnras/slx008>

P. Isola, et al., *Image-to-Image Translation with Conditional Adversarial Networks*, 2016 (Pix2Pix GAN)

[arXiv:1611.07004](https://arxiv.org/abs/1611.07004)

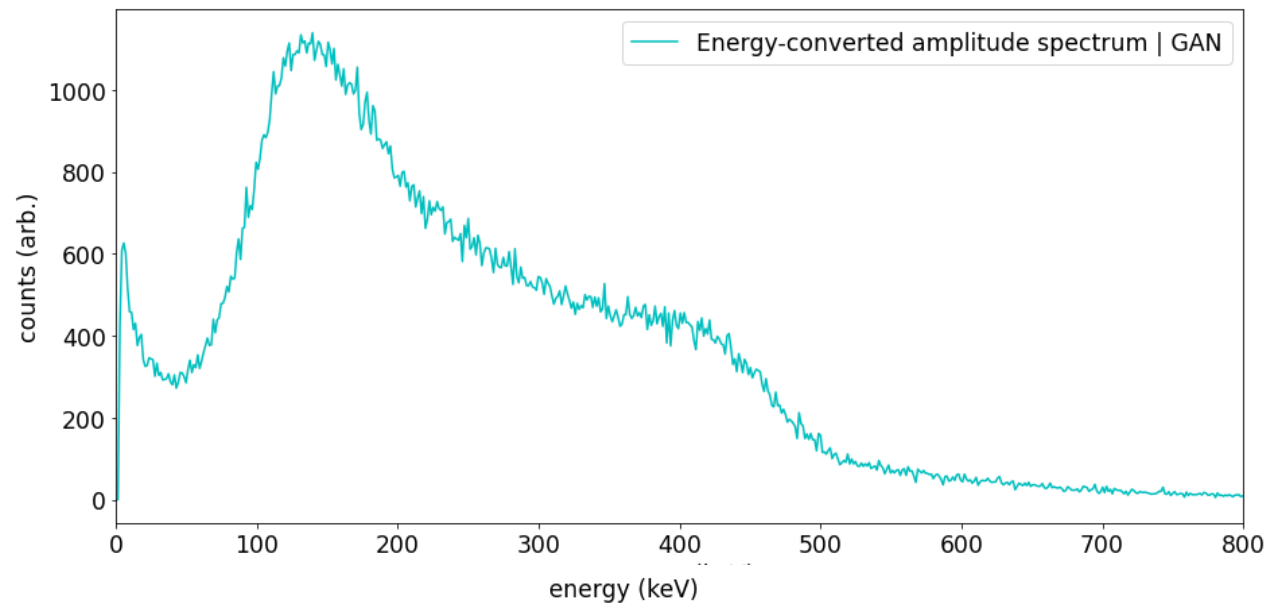
# GAN Predictions | LaBr3 Scintillator



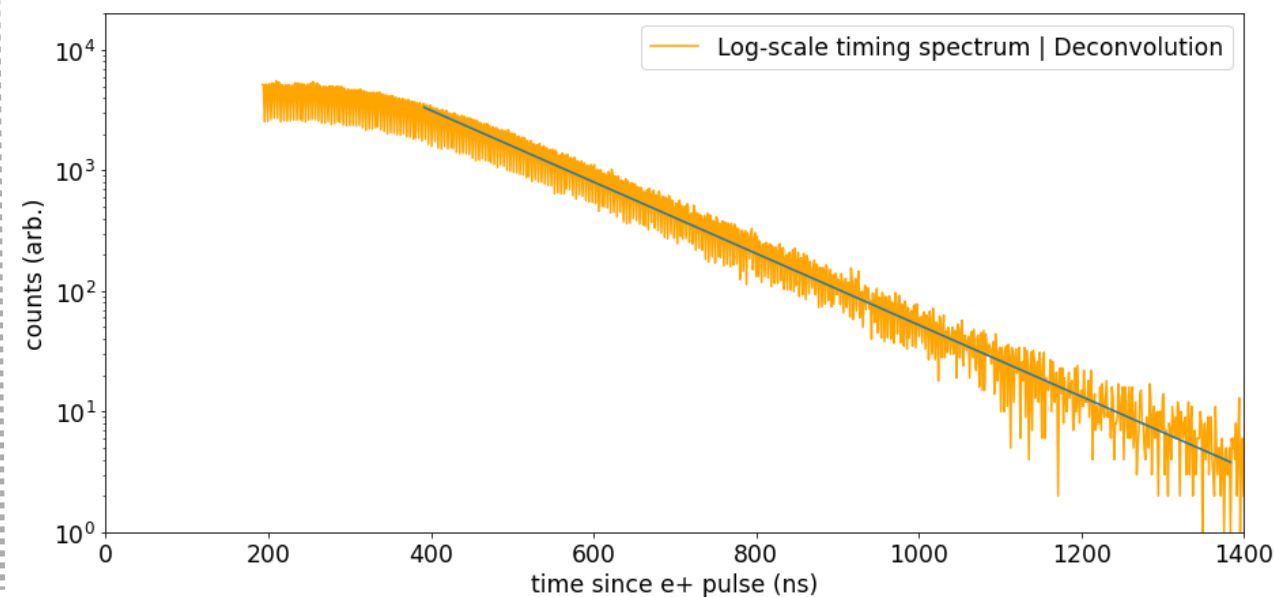
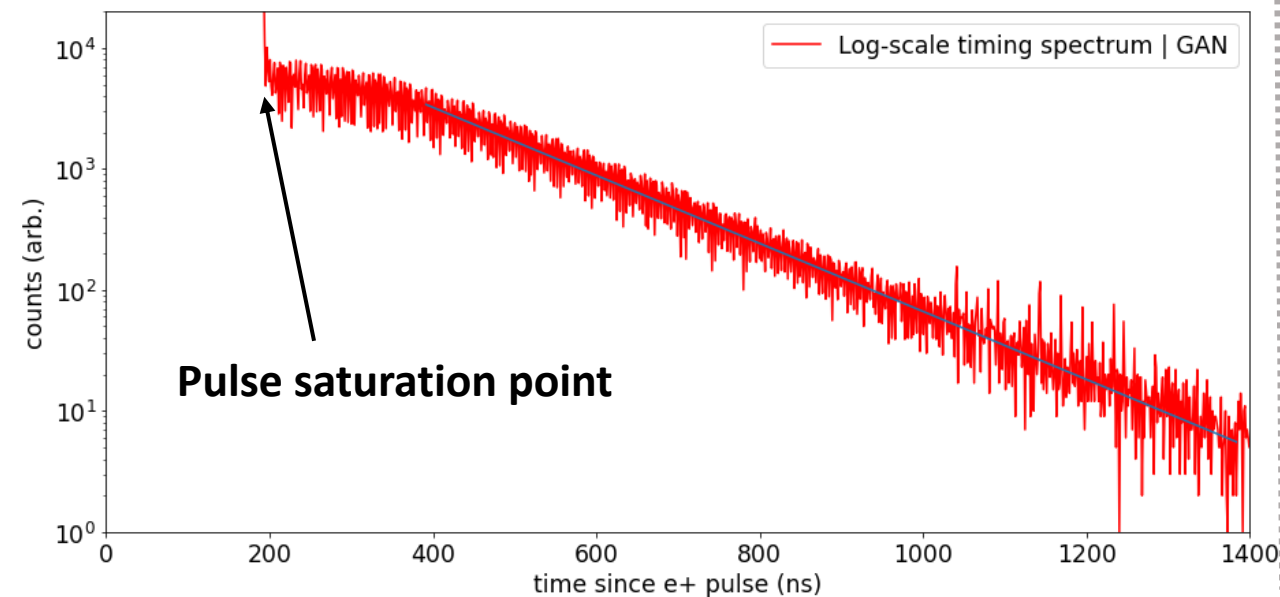
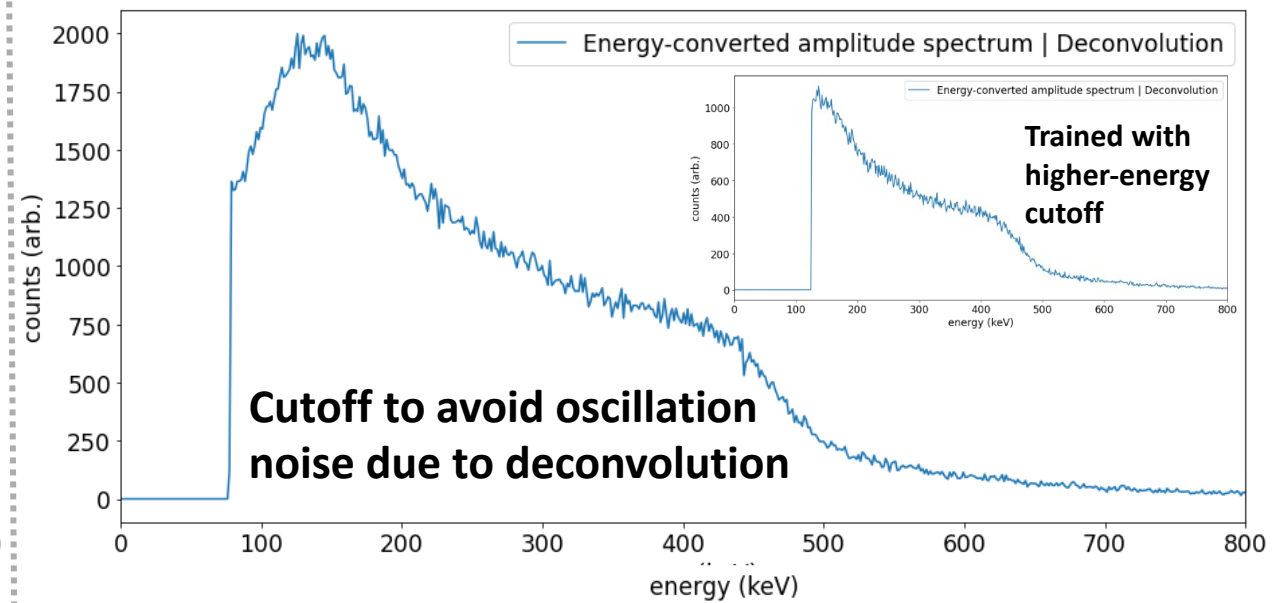
# Preliminary Results

Amplitudes are taken from 420 ns after e+ pulse due to incomplete deconvolution along tail (~190 ns for timing spectra)

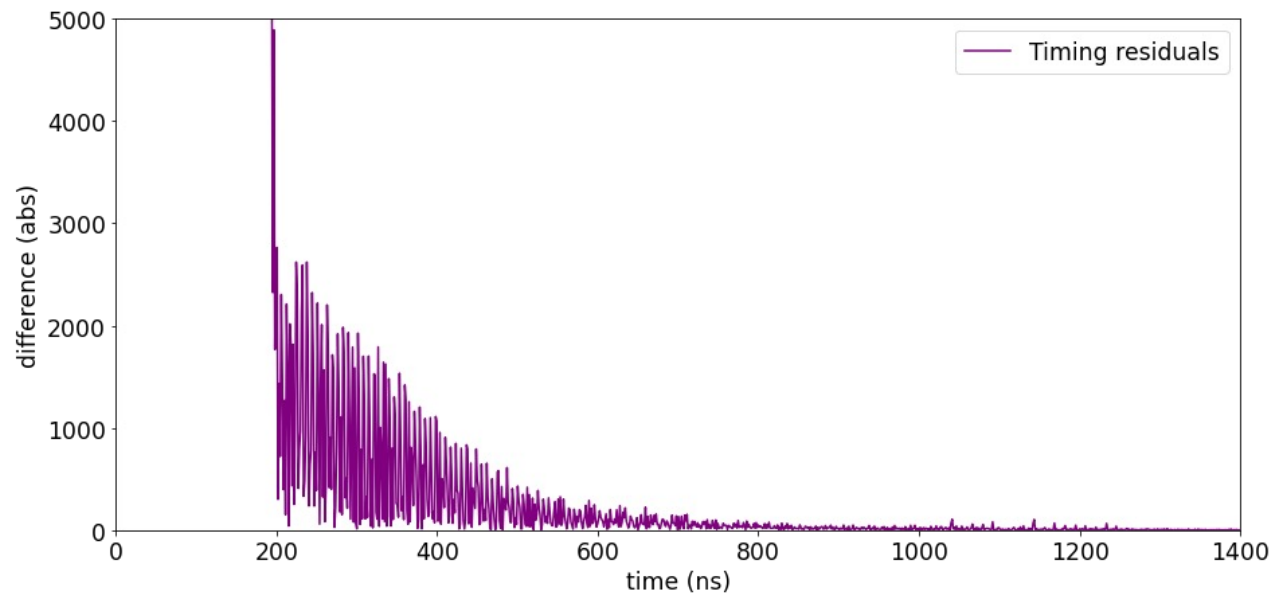
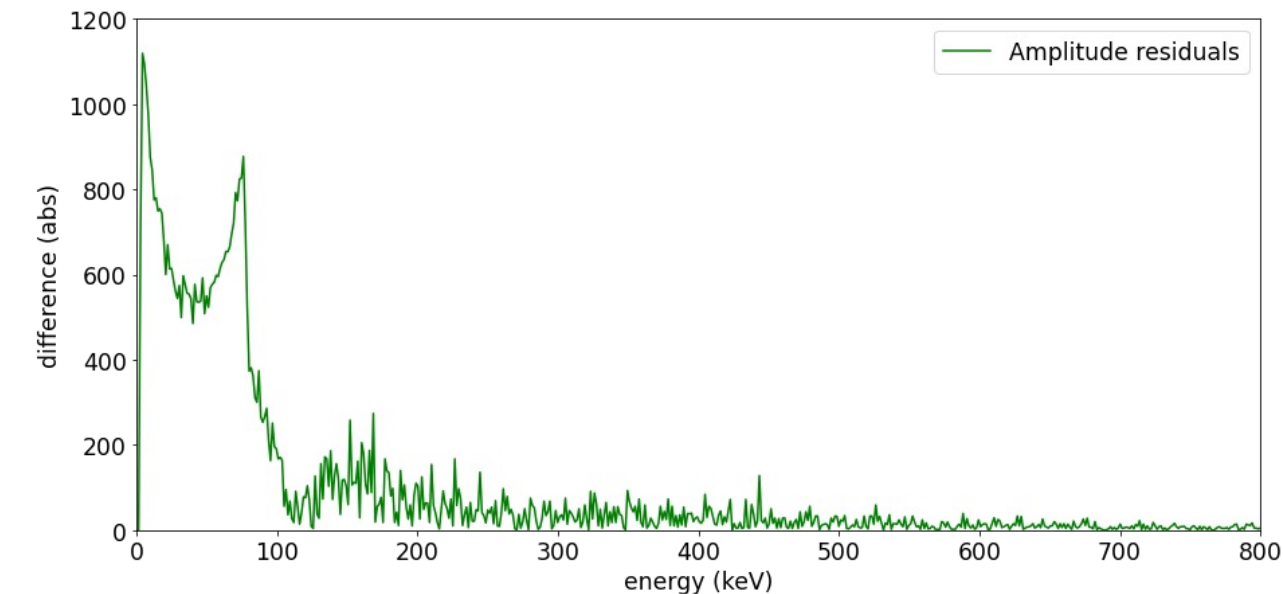
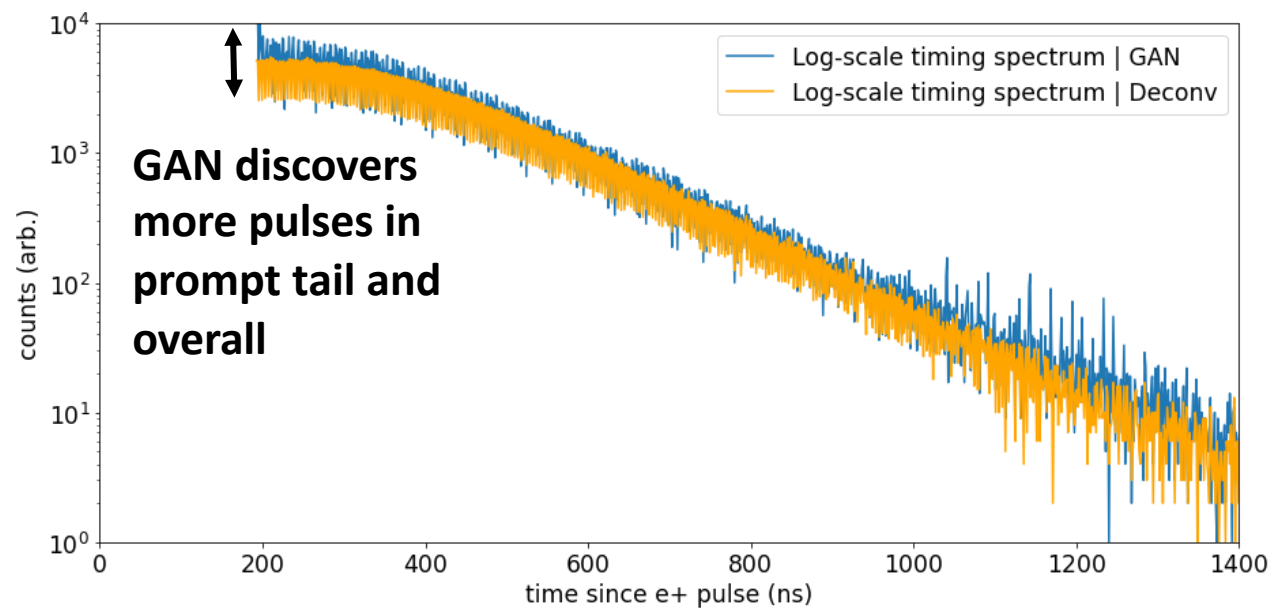
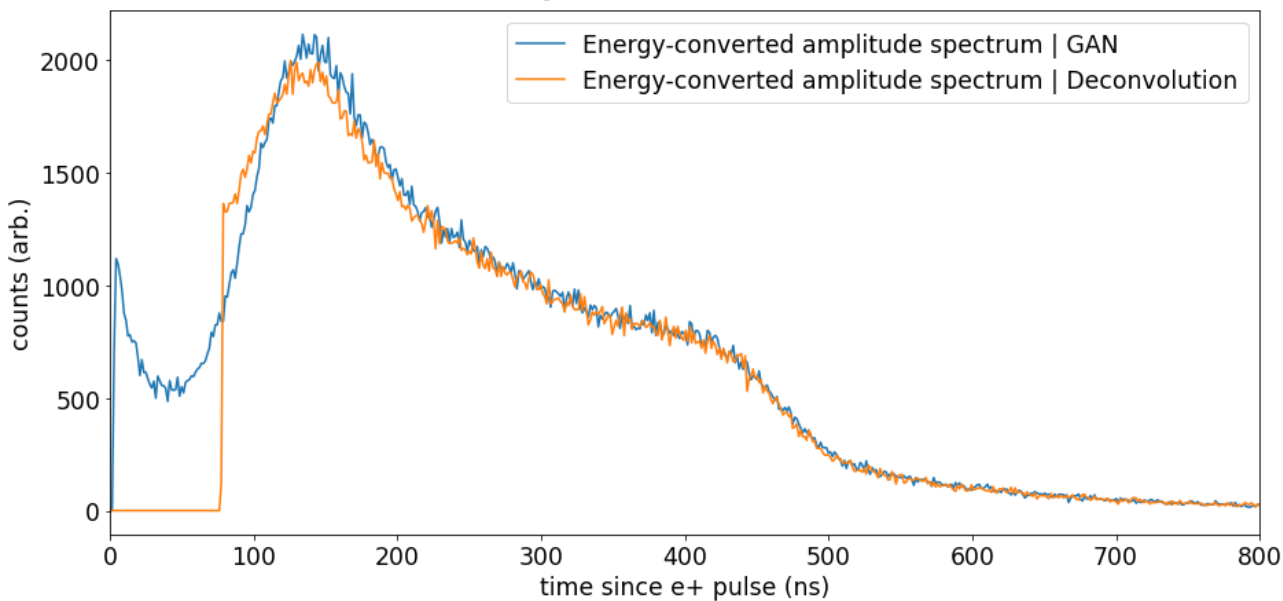
GAN



Deconvolution



# Preliminary Results

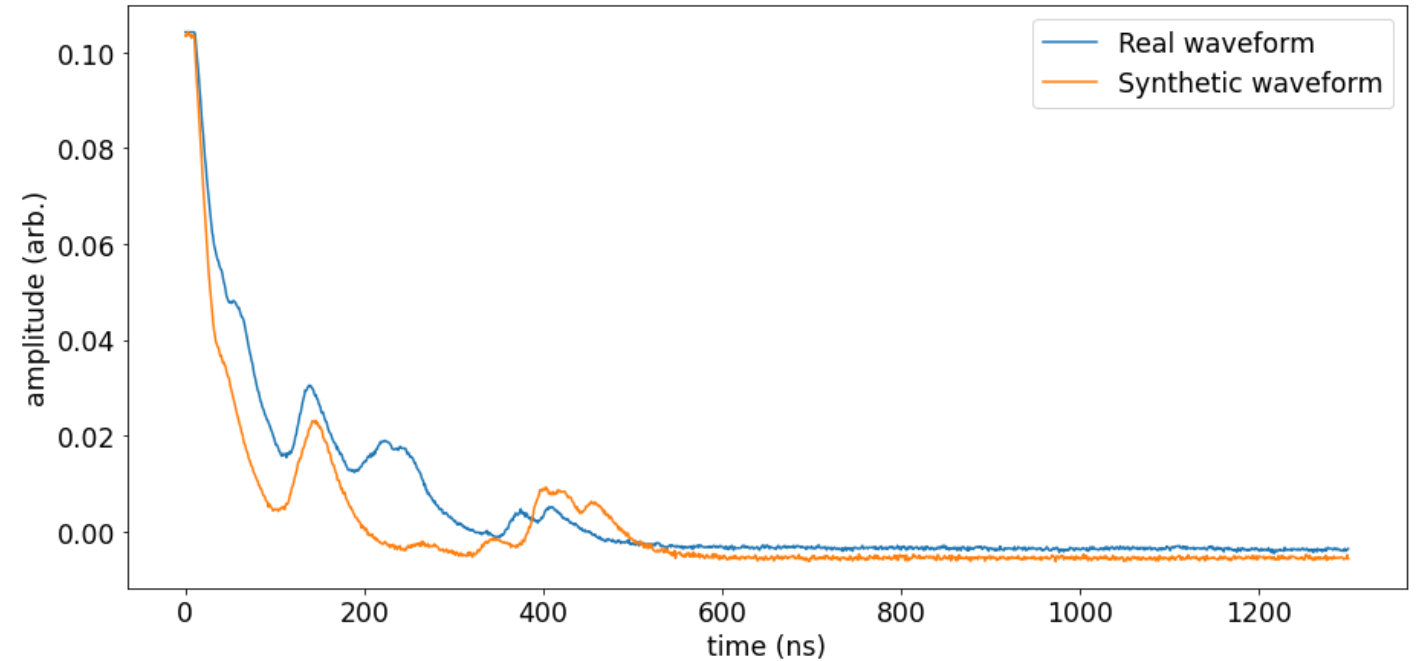


# Future

## Training on synthetic data

### Benefit:

- Know exact location and amplitude of peaks—i.e., not dependent on incomplete deconvolution of the prompt peak's exponential tail. This will potentially allow the full tail up to the pulse saturation level to be deconvoluted. The difficulty lies in make realistic synthetic data.



## Mixture of Experts

### Benefit:

--One model can deconvolute waveforms from a variety of detectors (LaBr3, NaI, plastic, HPGe, etc.)

